

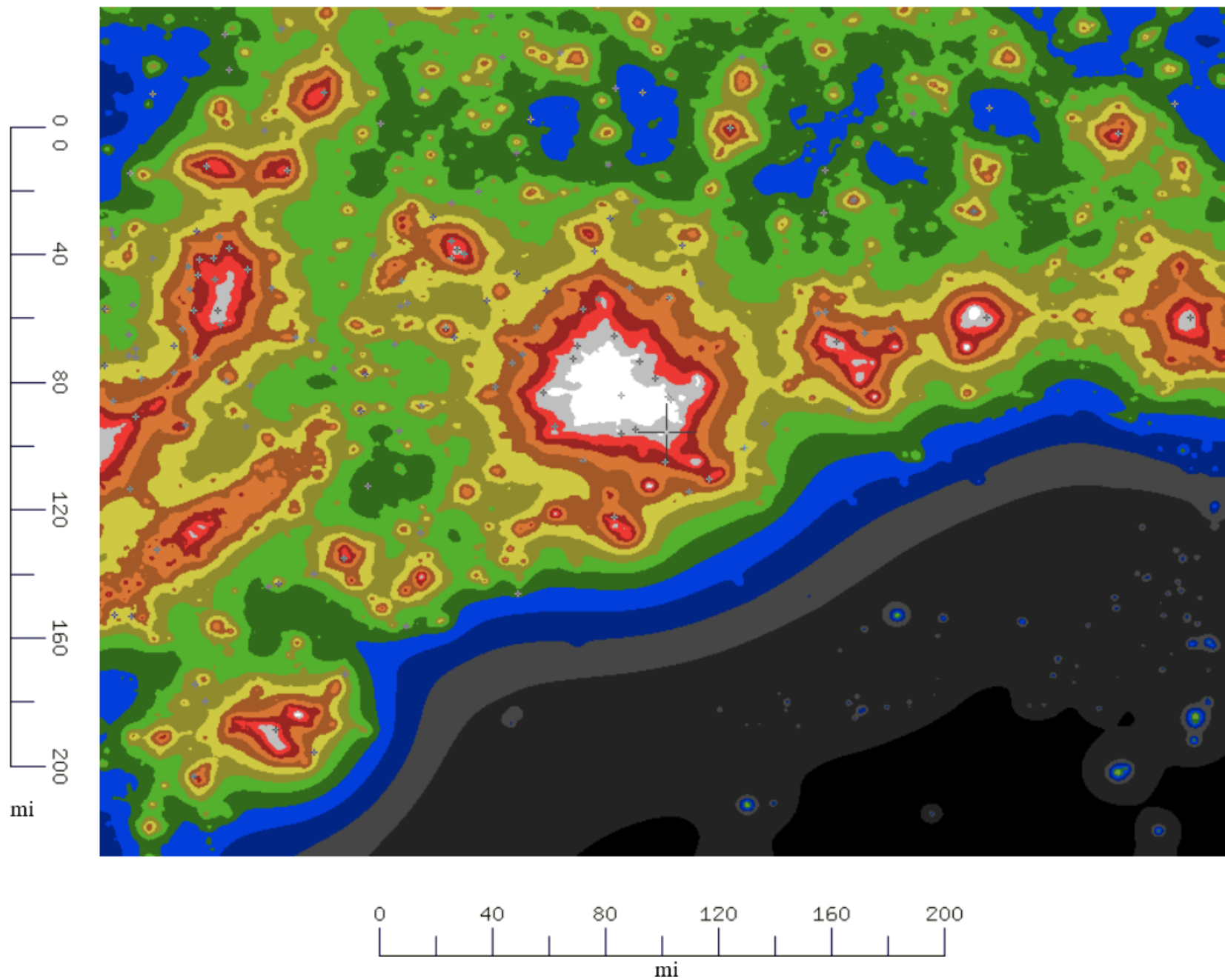
What Astronomy Can We Do from Here?

Doug Holland

What is the problem with doing astronomy from here?

Light Pollution -

Light Pollution Map of Our Area



The Bortle Scale

Conditions at Zenith				
Color	Artificial / Natural Sky Brightness	Sky Brightness mags / sq arcsec V Band	Bortle Scale <small>approx</small>	Description (Descriptions are approximate. Your sky may vary.)
	< 0.01	22.00 to 21.99	1	Theoretically darkest sky limited by airglow and starlight
	0.01 to 0.06	21.99 to 21.93	2	Gegenschein visible. Zodiacal light annoyingly bright. Rising milkyway confuses some into thinking it's dawn. Limiting magnitude 7.6 to 8.0 for people with exceptional vision. Users of large dobsonian telescopes are <i>very</i> happy. [-ad]
	0.06 to 0.11	21.93 to 21.89	2	Faint shadows cast by milkyway visible on white objects. Clouds are black holes in the sky. No light domes. The milky way has faint extentions making it 50 degrees thick. Limiting magntiude 7.1 to 7.5. [-ad]
	0.11 to 0.19	21.89 to 21.81	3	
	0.19 to 0.33	21.81 to 21.69	3	The sky is crowded with stars, extending to the horizon in all directions. In the absence of haze the M.W. can be seen to the horizon. Clouds appear as black silhouettes against the sky. Stars look large and close. [- Richard Berry] Low light domes (10 to 15 degrees) on horizon. M33 easy with averted vision. M15 is naked eye. Milky way shows bulge into Ophiuchus. Limiting magnitude 6.6 to 7.0. [-ad]
	0.33 to 0.58	21.69 to 21.51	4	21.6: ... a glow in the direction of one or more cities is seen on the horizon. Clouds are bright near the city glow. [- Richard Berry]
	0.58 to 1.00	21.51 to 21.25	4	Zodiacal light seen on best nights. Milkyway shows much dark lane structure with beginnings of faint bulge into Ophiuchus. M33 difficult even when above 50 degrees. Limiting magnitude about 6.2 to 6.5. [-ad]
	1.00 to 1.73	21.25 to 20.91	4.5	21.1: The M.W. is brilliant overhead but cannot be seen near the horizon. Clouds have a greyish glow at the zenith and appear bright in the direction of one or more prominent city glows. [- Richard Berry] Some dark lanes in milkyway but no bulge into Ophiuchus. Washed out milkyway visible near horizon. Zodiacal light very rare. Light domes up to 45 degrees. Limiting magnitude about 5.9 to 6.2. [-ad]
	1.73 to 3.00	20.91 to 20.49	4.5	
	3.00 to 5.20	20.49 to 20.02	5	20.4: To a city dweller the M.W. is magnificent, but contrast is markedly reduced, and delicate detail is lost. Limiting magnitude is noticeably reduced. Clouds are bright against the zenith sky. Stars no longer appear large and near. [- Richard Berry] Milkyway washed out at zenith and invisible at horizon. Many light domes. Clouds are brighter than sky. M31 easily visible. Limiting magnitude about 5.6 to 5.9.[-ad]
	5.20 to 9.00	20.02 to 19.50	5	
	9.00 to 15.59	19.50 to 18.95	6	19.5: M.W. is marginally visible, and only near the zenith. Sky is bright and discoloured near the horizon in the direction of cities. The sky looks dull grey. [- Richard Berry] Milkyway at best very faint at zenith. M31 difficult and indestinct. Sky is grey up to 35 degrees. Limiting magntidue 5.0 to 5.5. [-ad]
	15.59 to 27.00	18.95 to 18.38	7	
	27.0 to 46.77	18.38 to 17.80	8	Entire sky is grayish or brighter. Familliar constellations are missing stars. Fainter constellations are absent. Less than 20 stars visible over 30 degrees elevation in brigher areas. Limiting magntude from 3 to 4.CCD imaging is still possible. But telescopic visual observation is usually limited to the moon, planets, double stars and variable stars. [-ad]
	>46.77	>17.80	9	18.5: Stars are weak and washed out, and reduced to a few hundred. The sky is bright and discoloured everywhere. [- Richard Berry] Most people don't look up.[-ad]

We may wish we were on a mountain top in Chile...



But, there are things we can do from here =>

One thing we can do from here:

The Sun



<https://uamshealth.com/wp-content/uploads/2019/03/staring-at-the-sun.jpg>

Note – all astroimages are those of the author unless designated

The Sun

The Venus transit: 6/5/12

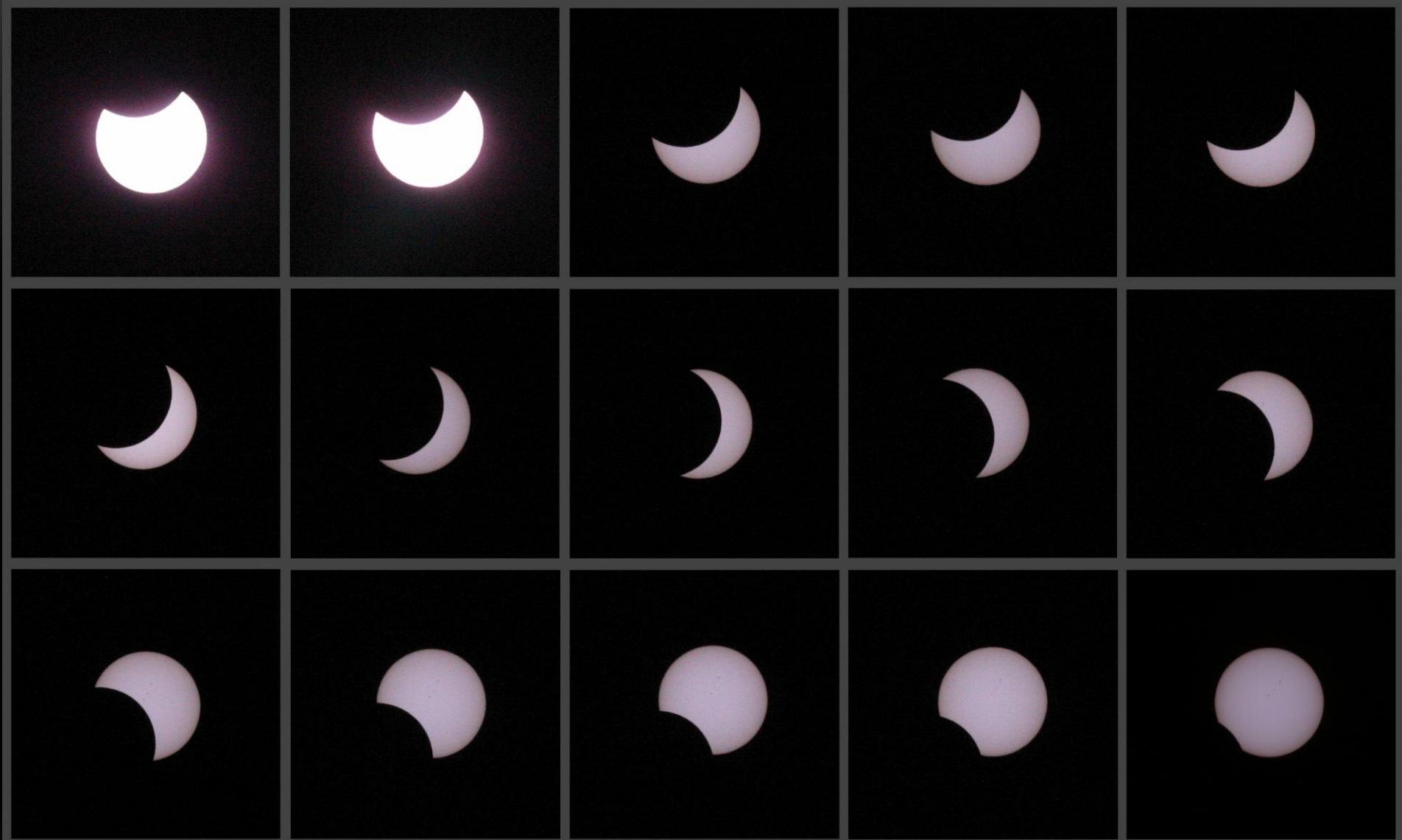


Mylar Solar Filter – visual & imaging



The Sun

Solar Eclipse:
8/21/17
from Houston



Solar Eclipse, 8/21/17

200mm Canon lens at f/22, 1/4000th Sec, Mylar Solar Filter, Canon 300D Camera

The Sun

Upcoming Eclipses:



Annular Solar Eclipse during ESP

Eldorado Star Party (ESP)

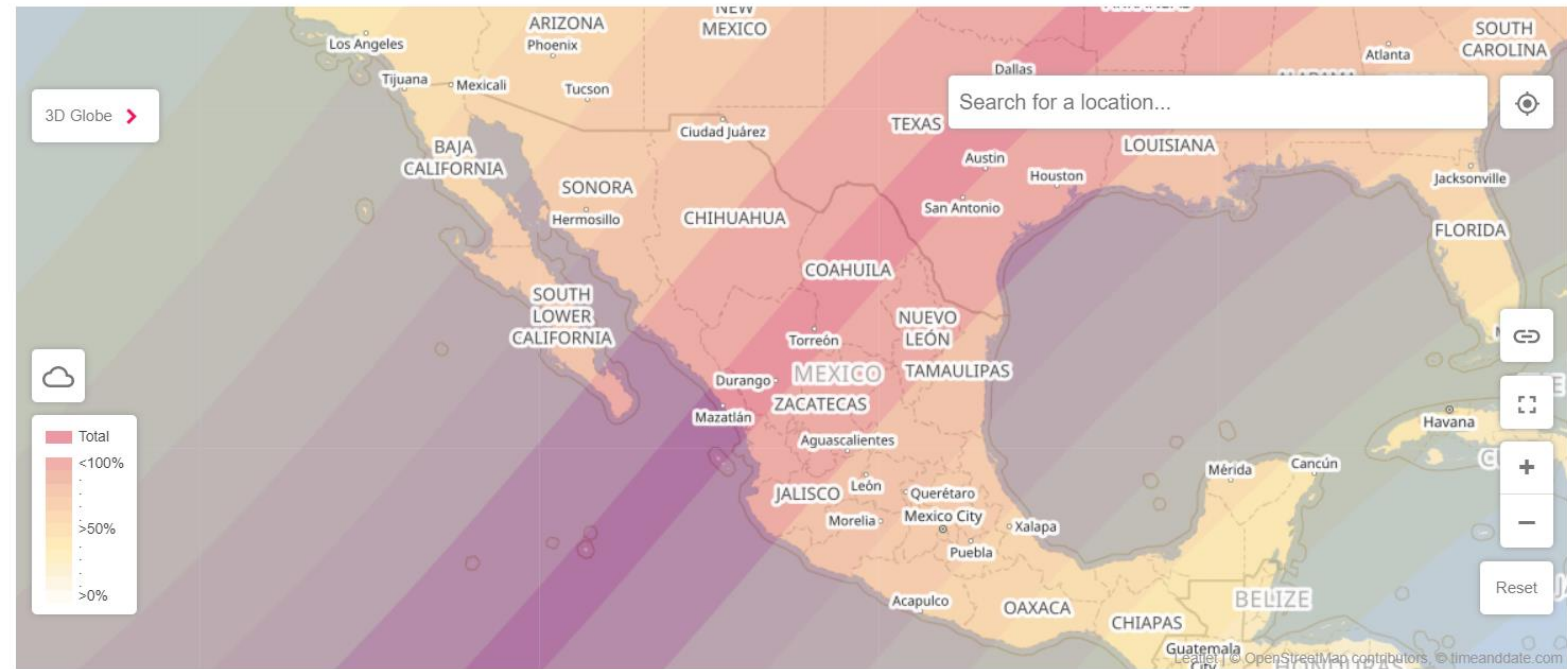
Saturday October 14, 2023

Peak 11:46 – 11:51 AM, over 5 minutes!



<https://preview.redd.it/mijj7gg4q2h01.jpg?auto=webp&s=7b826d279da2b30f6825f93a8f01a30fb6360d90>

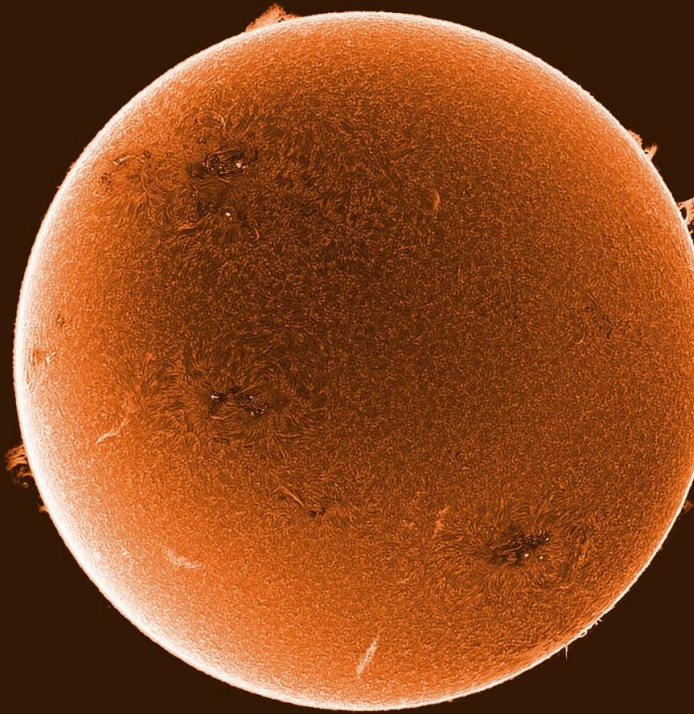
Eclipse Map — April 8, 2024 Total Solar Eclipse



The Sun

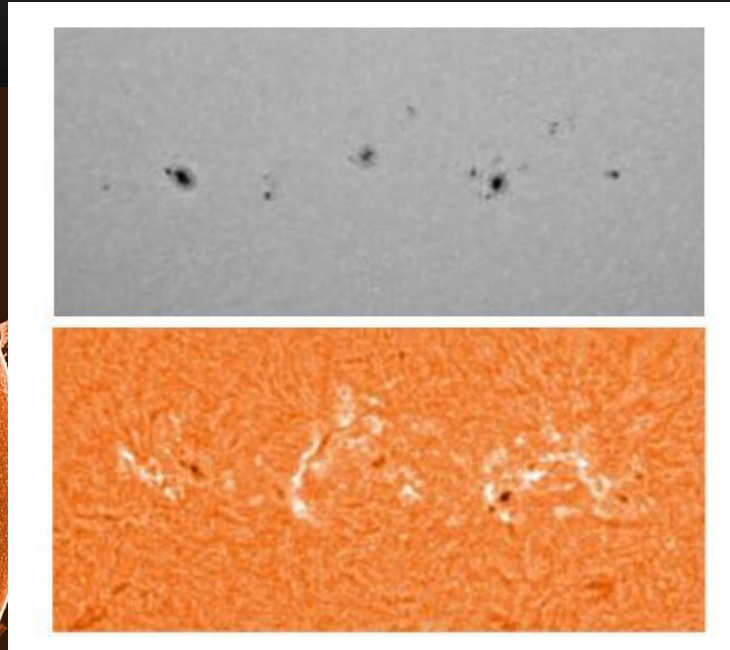
Solar Telescopes:

Solar sunspots and prominences

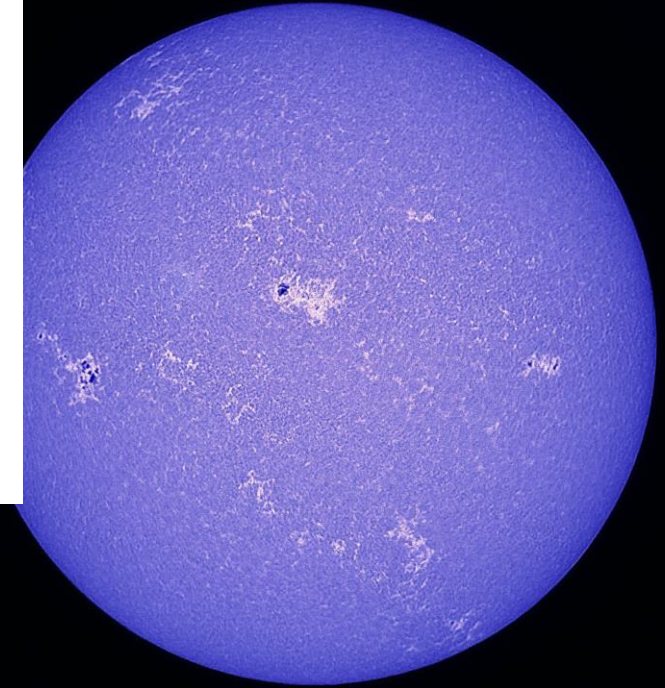


Hydrogen Alpha

656nm



**White light vs.
Hydrogen Alpha**
(granulation)



Calcium-K

393nm

Another thing we can do from here:

The Moon

1. Lunar Geography
2. Eclipses
3. High Res Imaging
4. Impacts



The Moon

2. Eclipses



Total Lunar Eclipse
December 20th - 21st, 2010
The Holland Observatory
200mm Newtonian at f/5
Canon Digital Rebel w/ MPCC



*Total Lunar Eclipse
January 20-21, 2019
The Holland Observatory
200mm Newtonian at f/5
Modified Canon 450D w/ MPCC*



1/4 Sec.

2 Sec.

8 Sec.

Lunar Eclipse - 5/15/22



10:21 PM



11:11 PM

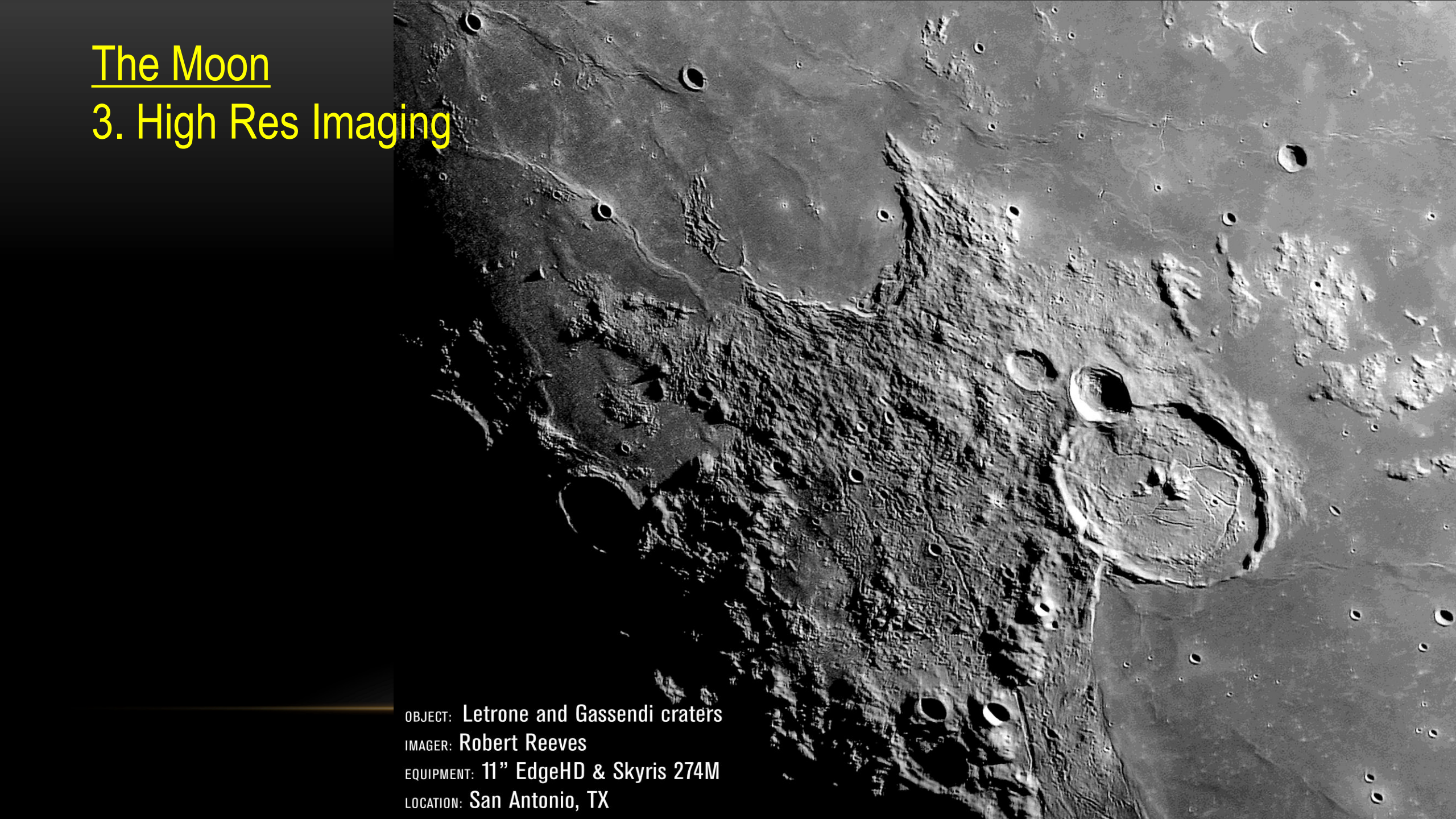


12:01 AM

*The Holland Observatory
80ED Refractor at f/7.5
AT2FF Field Flatteners
Modified Canon 450D Camera*

The Moon

3. High Res Imaging

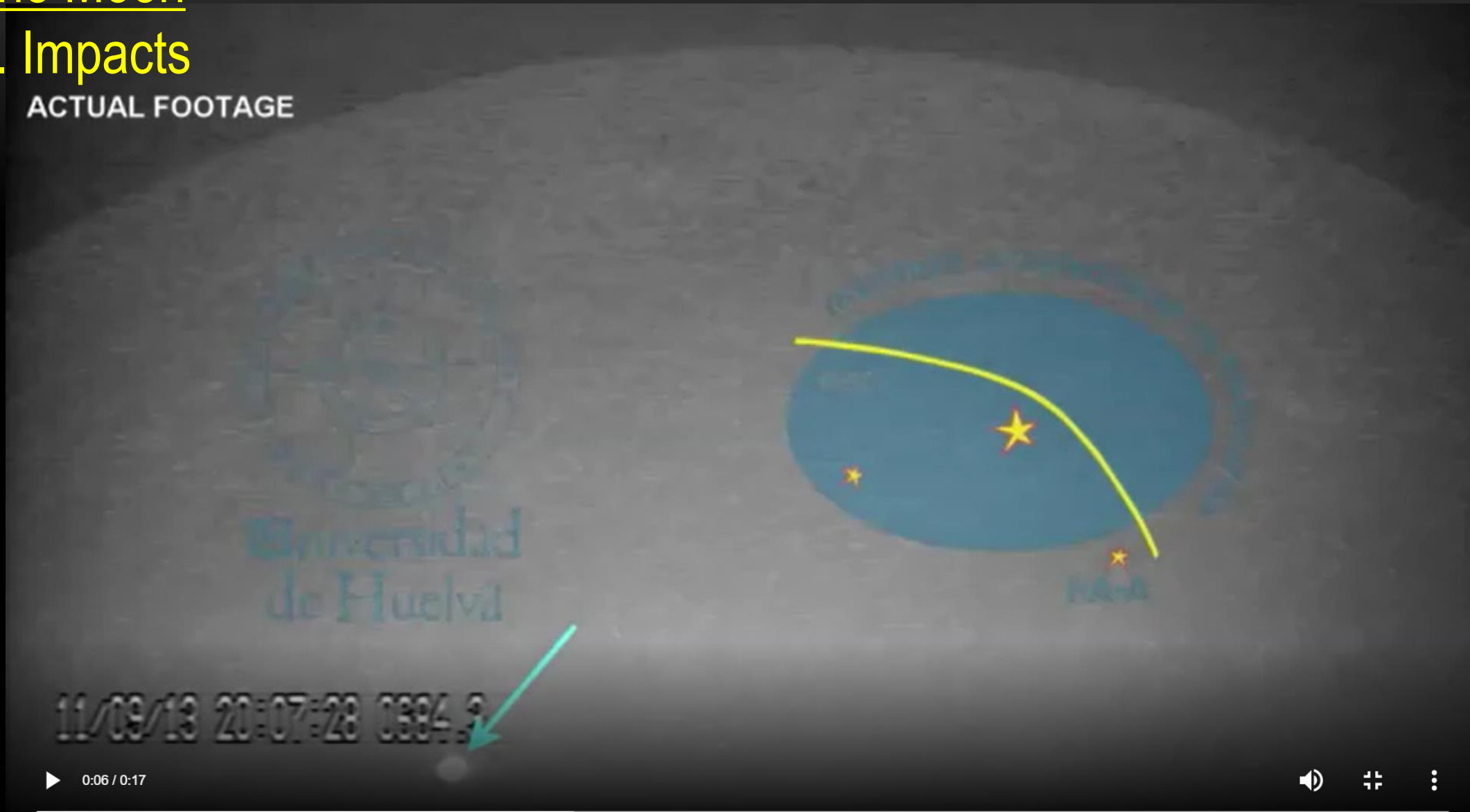
A high-resolution black and white photograph of the Moon's surface. The image shows a vast, cratered landscape. In the center-right, a large, prominent crater with a distinct rim and a dark interior is visible. To its left, another large crater is partially visible. The surface is covered with numerous smaller craters of various sizes, some with bright highlights and others in deep shadow. The lighting creates strong shadows, emphasizing the rugged terrain and the depth of the craters. The overall appearance is that of a heavily cratered, ancient celestial body.

OBJECT: Letrone and Gassendi craters
IMAGER: Robert Reeves
EQUIPMENT: 11" EdgeHD & Skyris 274M
LOCATION: San Antonio, TX

The Moon

4. Impacts

ACTUAL FOOTAGE



What else can we do?

Our Solar System Planets

1. Visual observing
 2. Imaging
 3. Phenomena of solar system planets
-

Our Solar System Planets

1. Visual observing

Even with small telescope:

- All planets visible
- Rings of Saturn
- Moons of Jupiter & Saturn
- High brightness – able to see in light polluted areas:
Haak Winery Uranus example



Our Solar System Planets

2. Imaging

- Planetary Camera
- Hundreds of images
- Stacked and processed

*** In general – seeing is better here than other places in country, including West Texas !! Good seeing is key to planetary imaging.

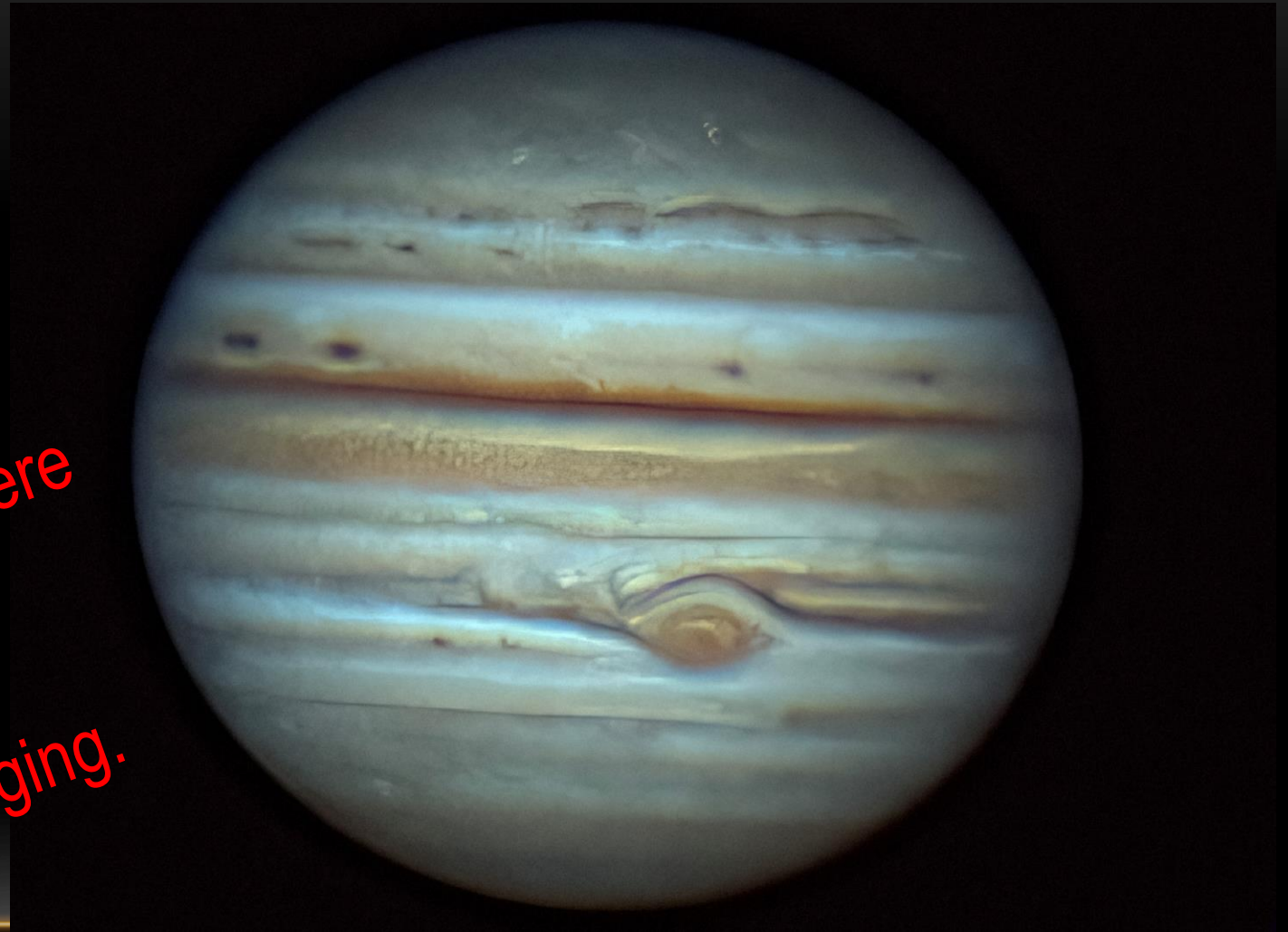


Image produced by Treavor Quinn

Our Solar System Planets

3. Phenomena of solar system planets

Many examples – a few here:

- Tilt of Saturn's rings
- Movement of Jupiter's Moons
 - Able to see IO disappear and reappear in one night
- Venus phases, similar to Moon
- Ice caps on Mars, changing with season
- Impacts on Jupiter

What about other solar system objects?

Meteor showers, comets, asteroids

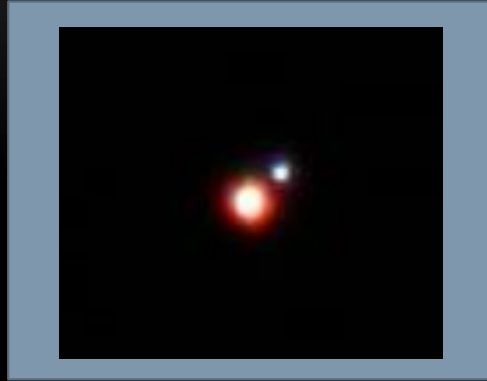
=> Can be viewed & imaged but generally much better in dark sky areas



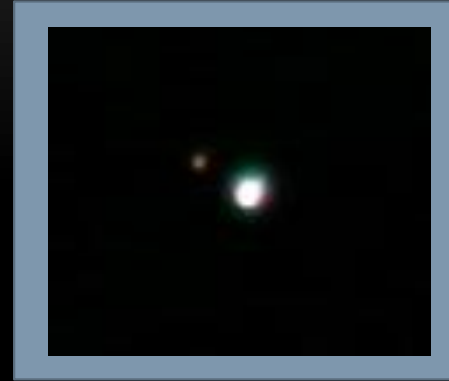
What else can we do?

Moving outside our Solar System =>

Multiple Star Systems (double stars)



Almach
in Andromeda



Eta Cassiopeiae

- Most star systems in the Milky Way are multiple star systems
- One celestial object that exhibits color with visual observing
- Many in view throughout the year
 - UHCL Star Party 11/16/22

Globular Clusters



- Able to see visually
- Image with short (1 to 2 minute) exposures – avoids being overwhelmed by light pollution

Open Clusters



The Holland Observatory
1/30/15

M37 - Open Cluster in Auriga
LRGB - L: 12x0.5min, R: 6x1min, G: 6x1min, B: 6x1min
AT2PC 1/8 AT2F5 SC8300

- Able to see visually & image
- Image with short (0.5 to 2 minute) exposures – avoids being overwhelmed by light pollution

But ---



The Holland Observatory
Fort McKavett
10/23/14

M45 - The Pleiades in Taurus
80ED f/7.5, Canon 450D 29x5min

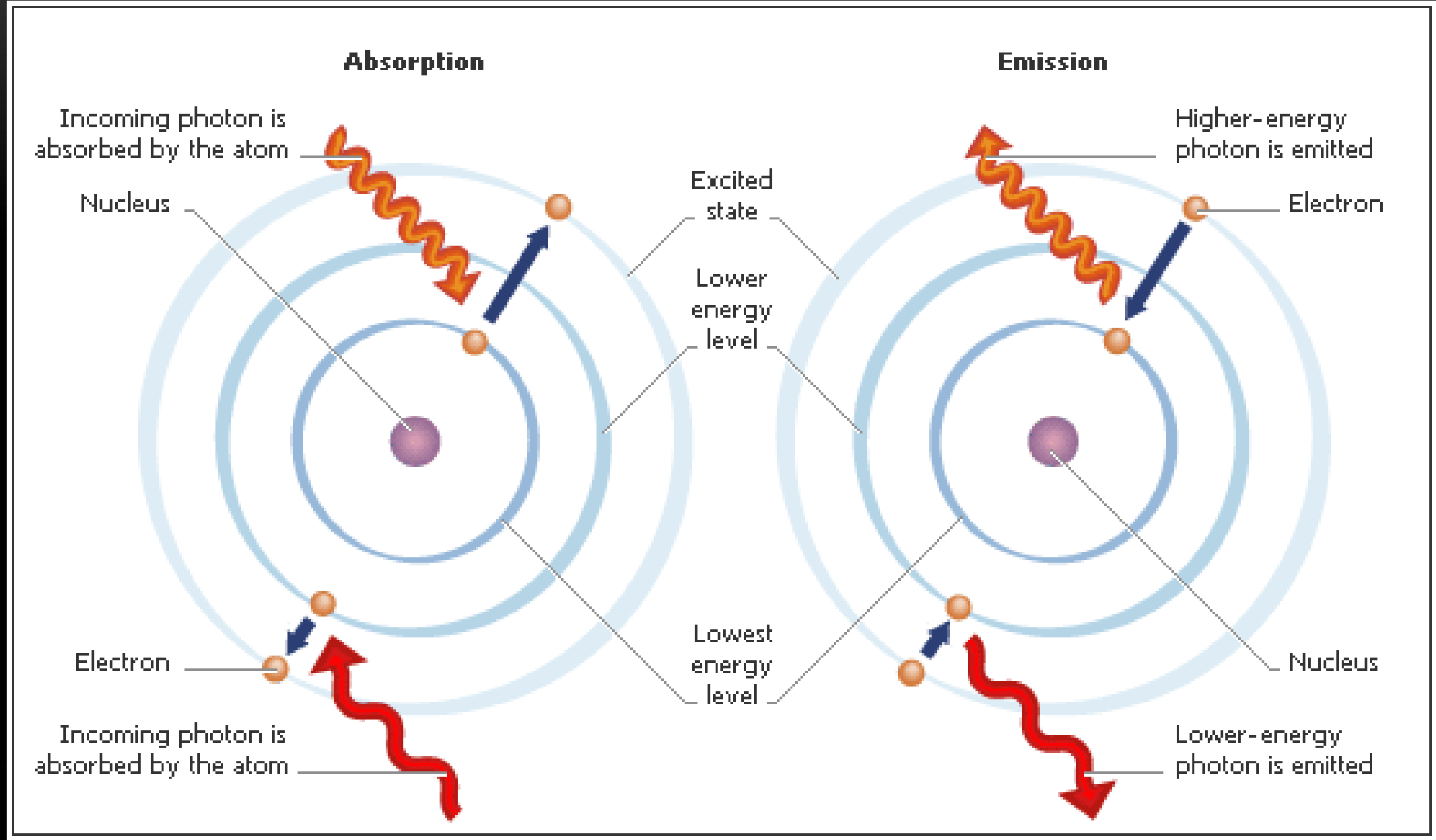
Deep images of Open Clusters with nebulosity requires a dark sky...

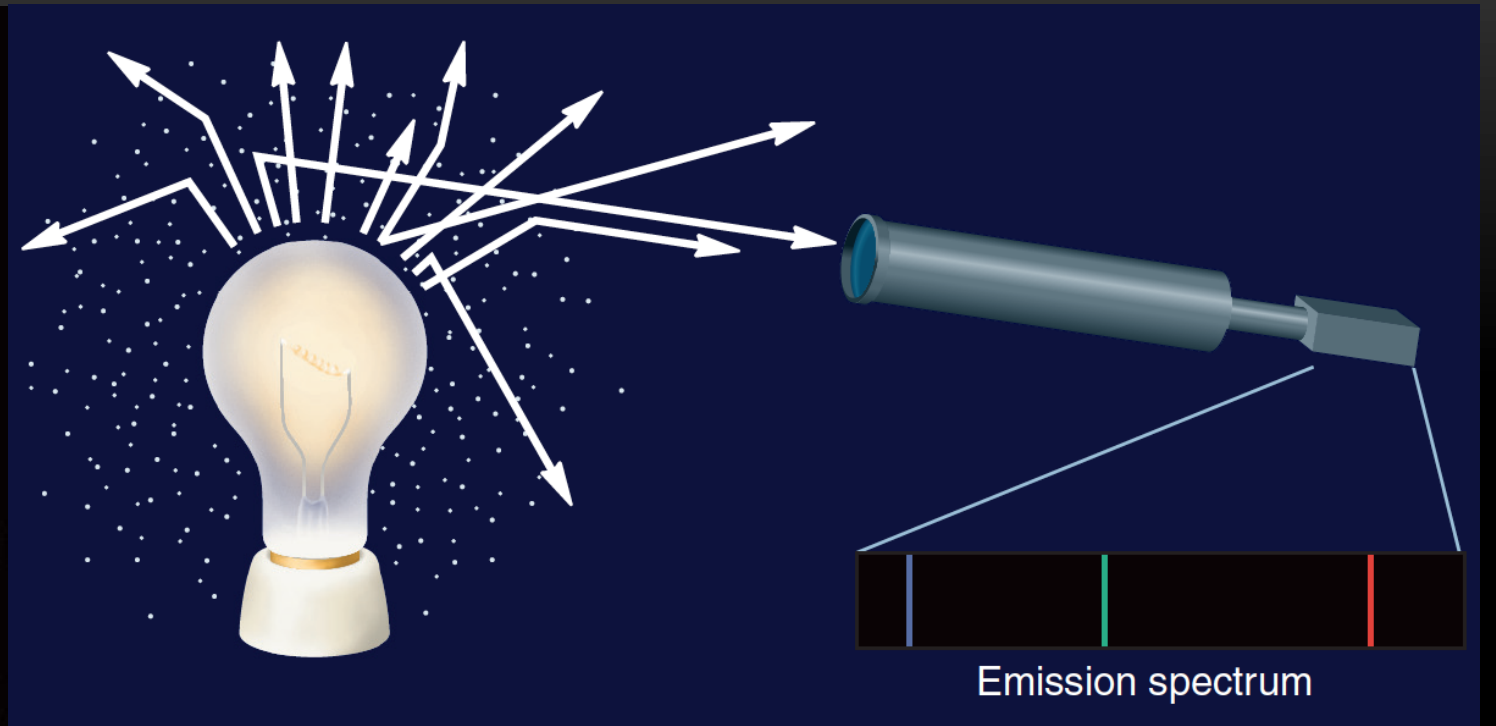
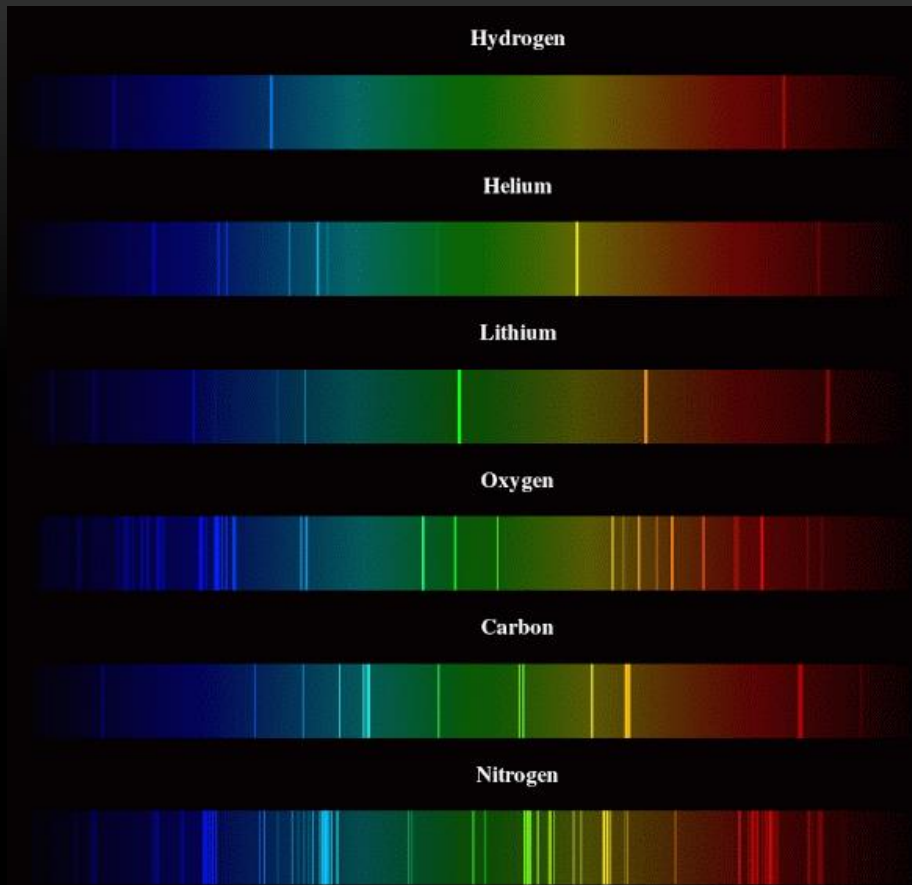
What else can we do?

Emission Nebulae



What is an Emission Nebula?





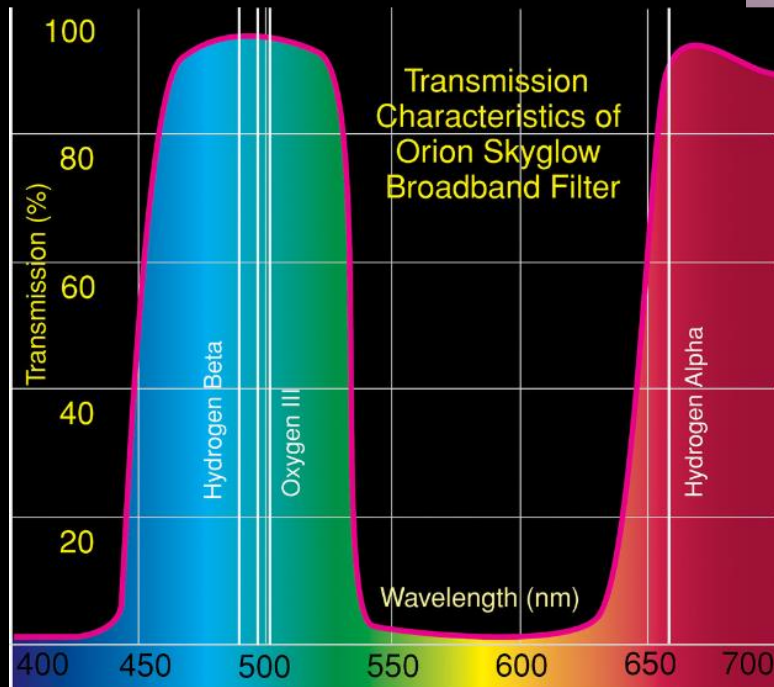
How can Emission Nebula be imaged in light polluted areas?

Since Emission Nebula emit at specific wavelengths => Filters



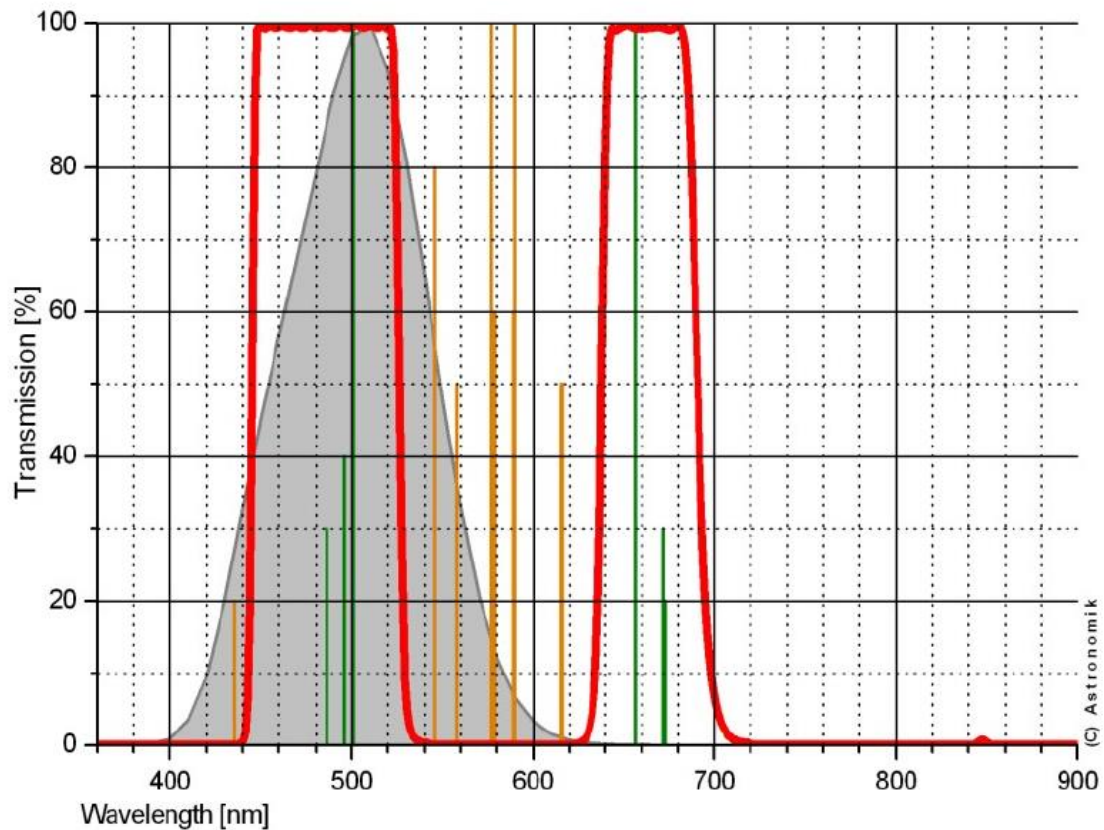
Orion Nebula with and without Skyglow Filter

<= Pass emission wavelengths, block others

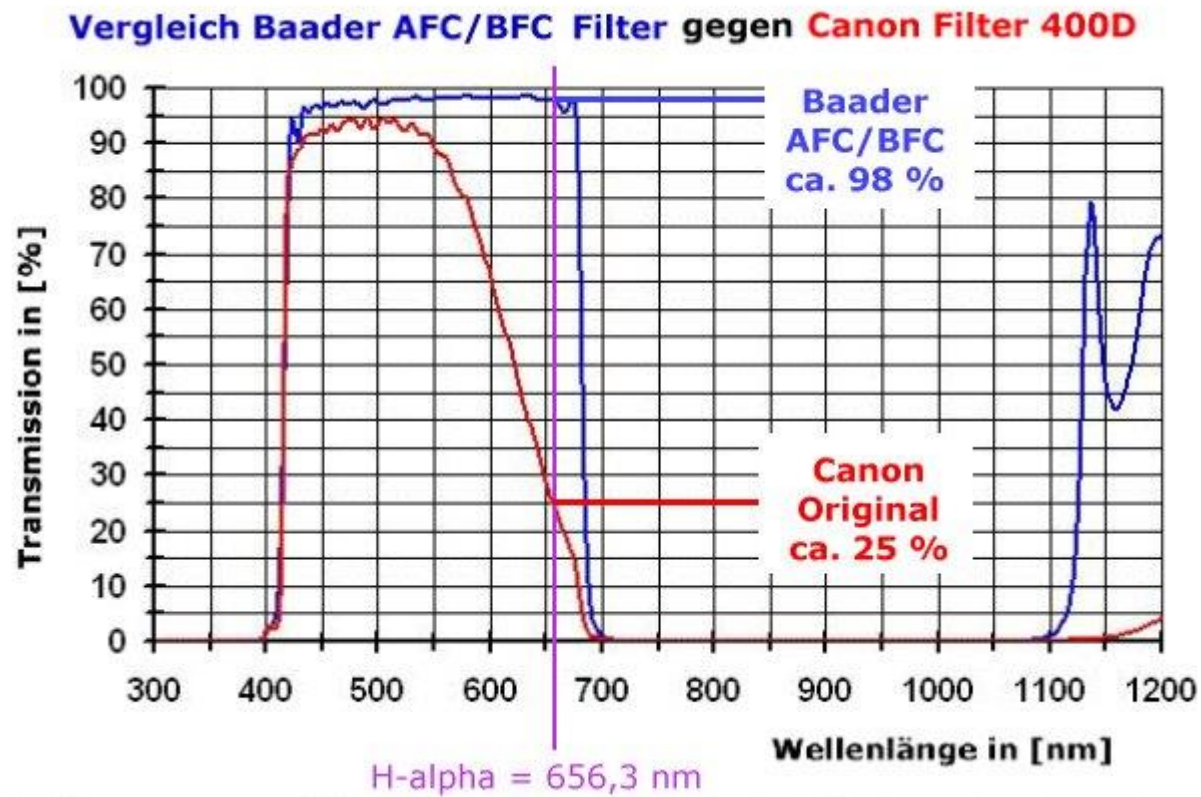




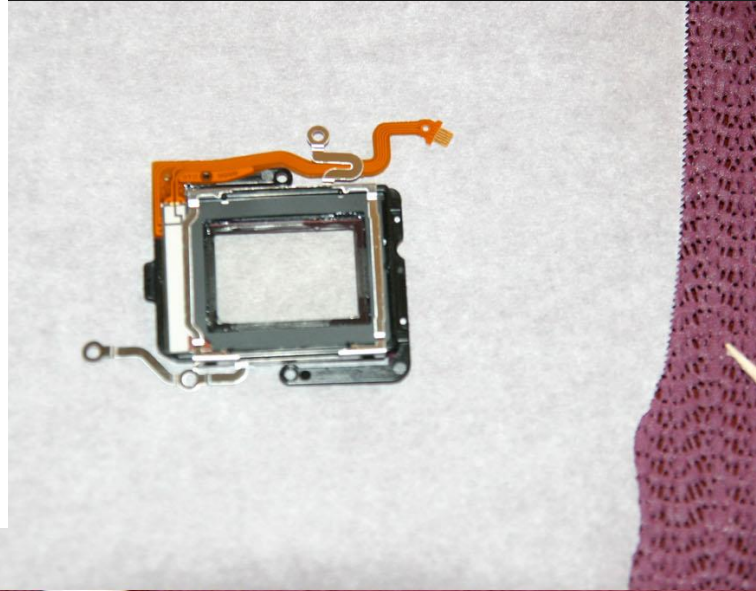
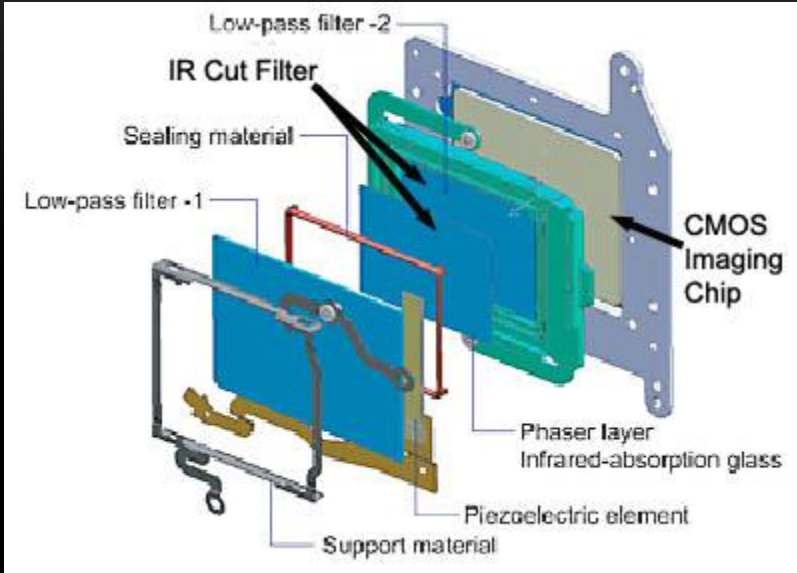
Previous image taken with Astronomik CLS Filter & Modified Canon DSLR



The major emission lines of artificial light pollution: The major emission lines of nebulae:



A quick detour => Modified Canon DSLR??



Replace IR Cut & Low Pass Filter #2 with new Baader UV / IR Filter



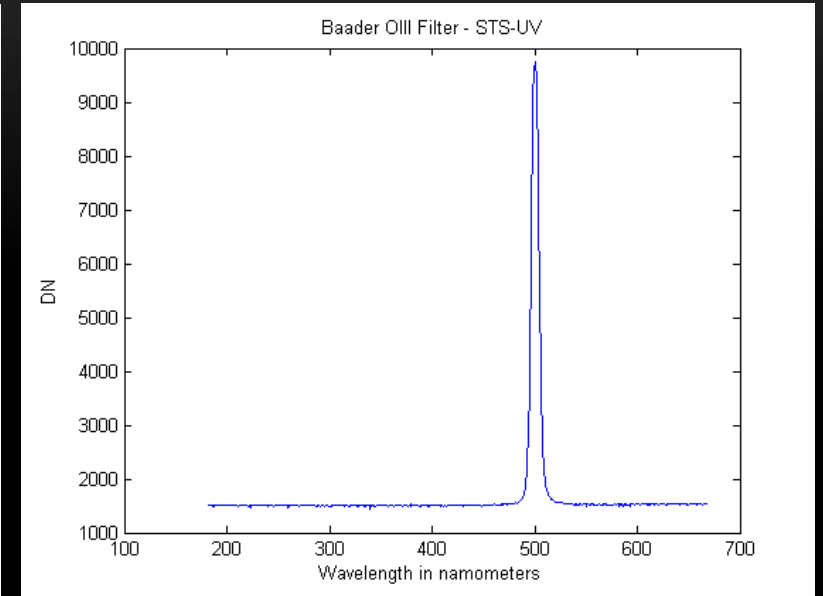
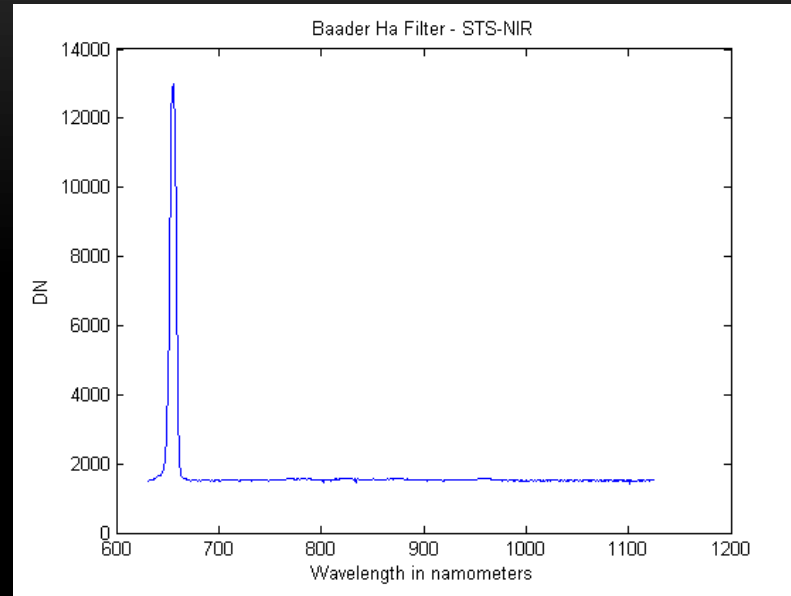
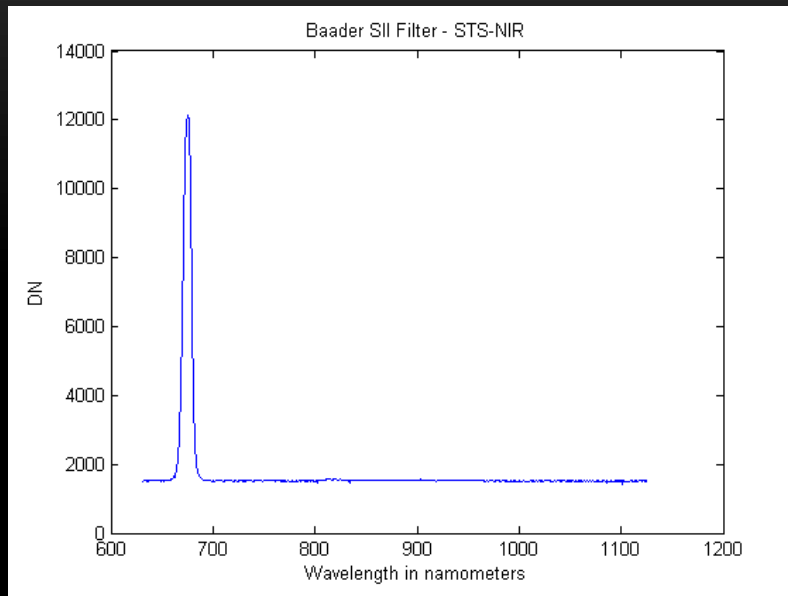
What else can we do?

Narrowband Imaging

Note - Previous filters are considered “broadband” filters

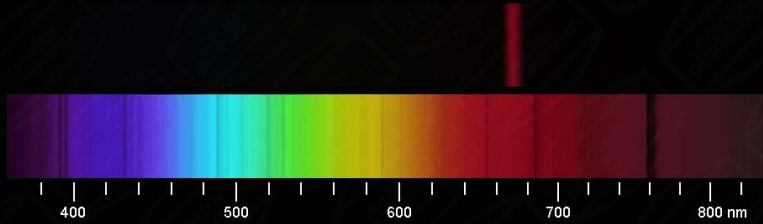
Generally done with dedicated astronomical camera, rather than DSLR





Baader S II-CCD 2458432 Narrowband-Filter 8nm
Real Filter-Spectrum, taken with MPI-DADOS-Spectrograph

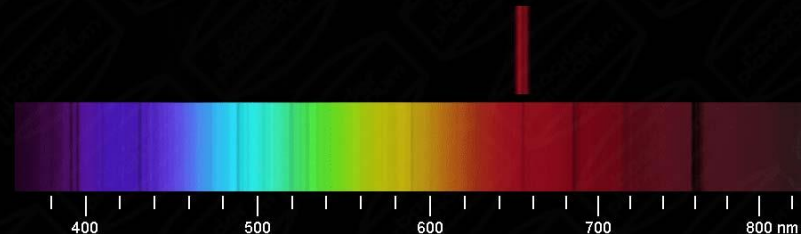
Ca Hg H β OIII Hg Na H α O₂ atm. O₂ atm.



Comparison Solar-Spectrum

Baader 7nm H-alpha CCD 2458383 Filter
Real Filter-Spectrum, taken with MPI-DADOS-Spectrograph

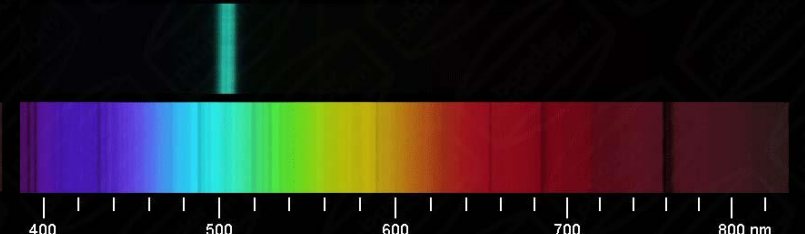
Ca Hg H β OIII Hg Na H α O₂ atm. O₂ atm.



Comparison Solar-Spectrum

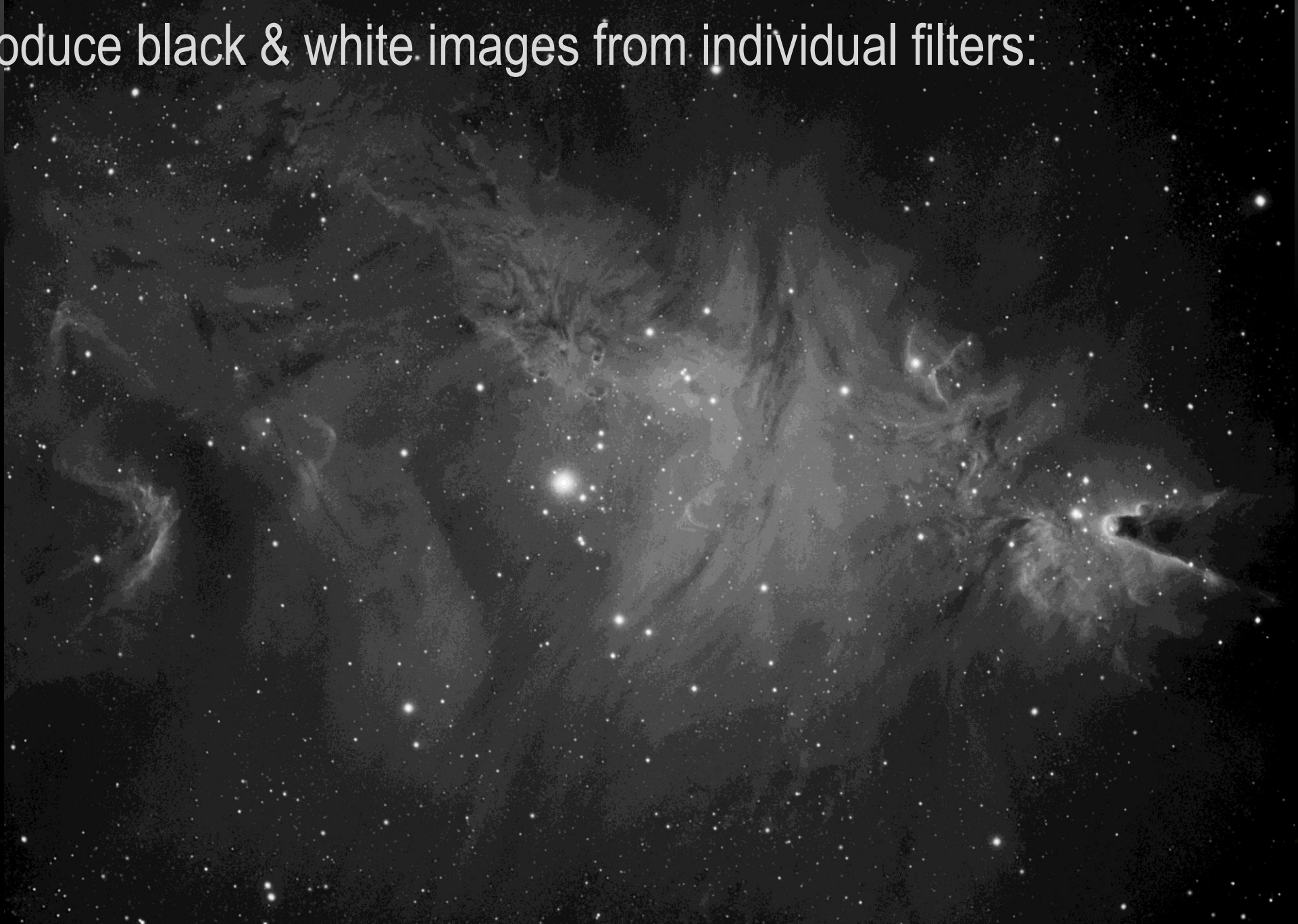
Baader OIII-CCD 2458437 Narrowband-Filter 8,5nm
Real Filter-Spectrum, taken with MPI-DADOS-Spectrograph

Ca Hg H β OIII Hg Na H α O₂ atm. O₂ atm.



Comparison Solar-Spectrum

Can produce black & white images from individual filters:





The Holland Observatory
5/5-6/21

M16 - The Eagle Nebula in Serpens
Pseudo Luminance - Ha: 18x5min + OIII: 24x5min
AT8RC at f/8, SC285 Camera





SII
29x15min



Ha
31x15min



OIII
33x15min

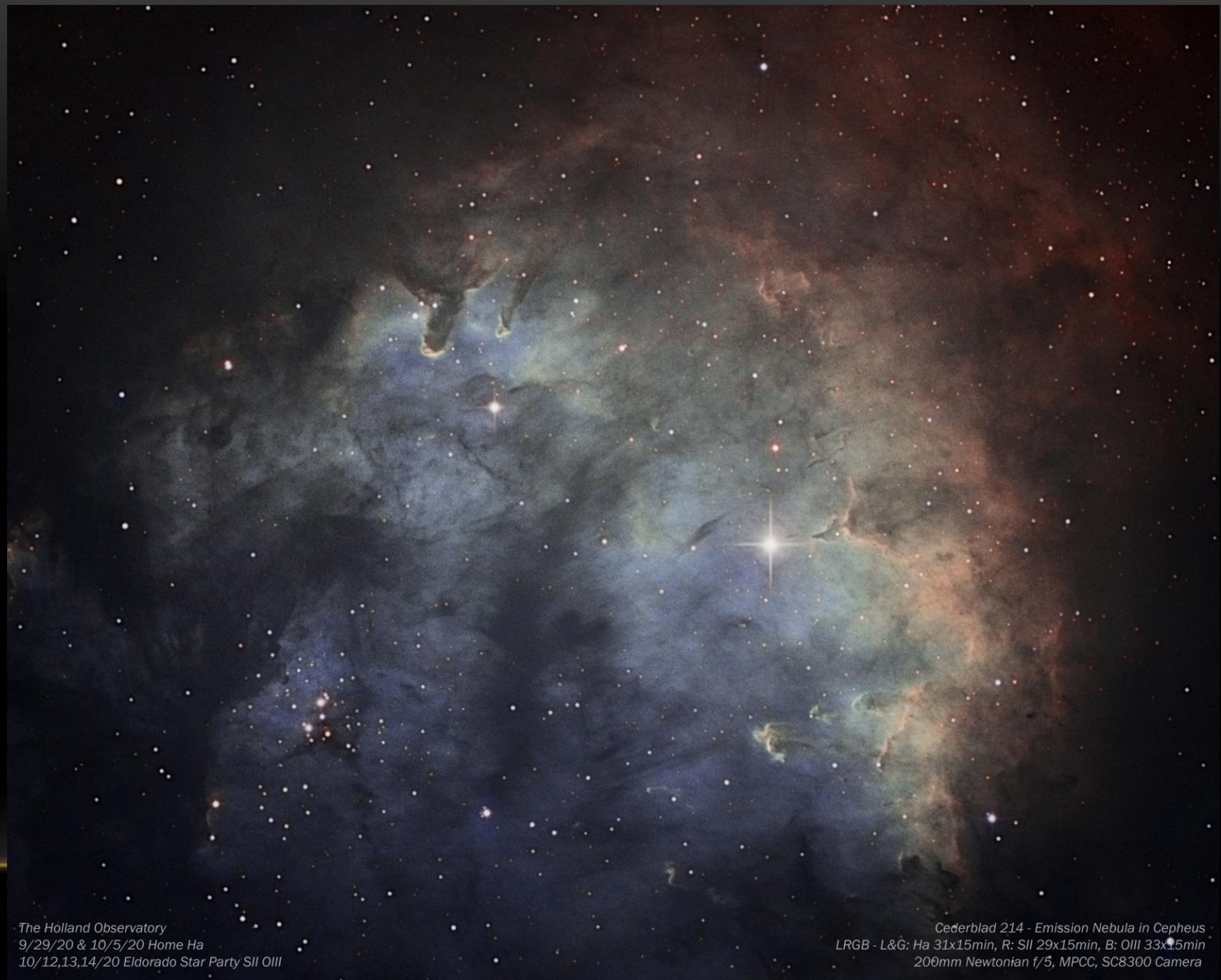
Total 93x15min => 23hours, 15 minutes

Narrowband Imaging

Can produce color images from individual filters by assigning filters to different colors

Processing

- Calibration: DSS (Deep Sky Stacker) [Provided better results than PixInsight (PI)]
- PI Dynamic Crop
- PI MultiScale Linear Transform (MLT) noise reduction for SII & OIII
- Photoshop (PS) Levels & Curves, Ha Smart Sharpen with star mask
- PS LRGB combine
- PI Correct Magenta Stars script
- PS Star blur w/ star mask
- PS Reduce Color Noise
- PS Dust & Scratches
- PI Subtractive Color Noise Reduction (SCNR) 87% + Color Mask script 13% to fix colors



The Holland Observatory
9/29/20 & 10/5/20 Home Ha
10/12,13,14/20 Eldorado Star Party SII OIII

Cederblad 214 - Emission Nebula in Cepheus
LRGB - L&G: Ha 31x15min, R: SII 29x15min, B: OIII 33x15min
200mm Newtonian f/5, MPCC, SC8300 Camera



The Holland Observatory
1/27 - 28/12

NGC2244, 2237 - 2239, The Rosette Nebula in Monoceros
L(Ha), R(SII), G(Ha), B(OIII - Ha): 20x3min, SII: 13x3min + 40x4min; OIII: 20x3min
200mm f/2.8 Canon L Series Lens, SC285 Camera



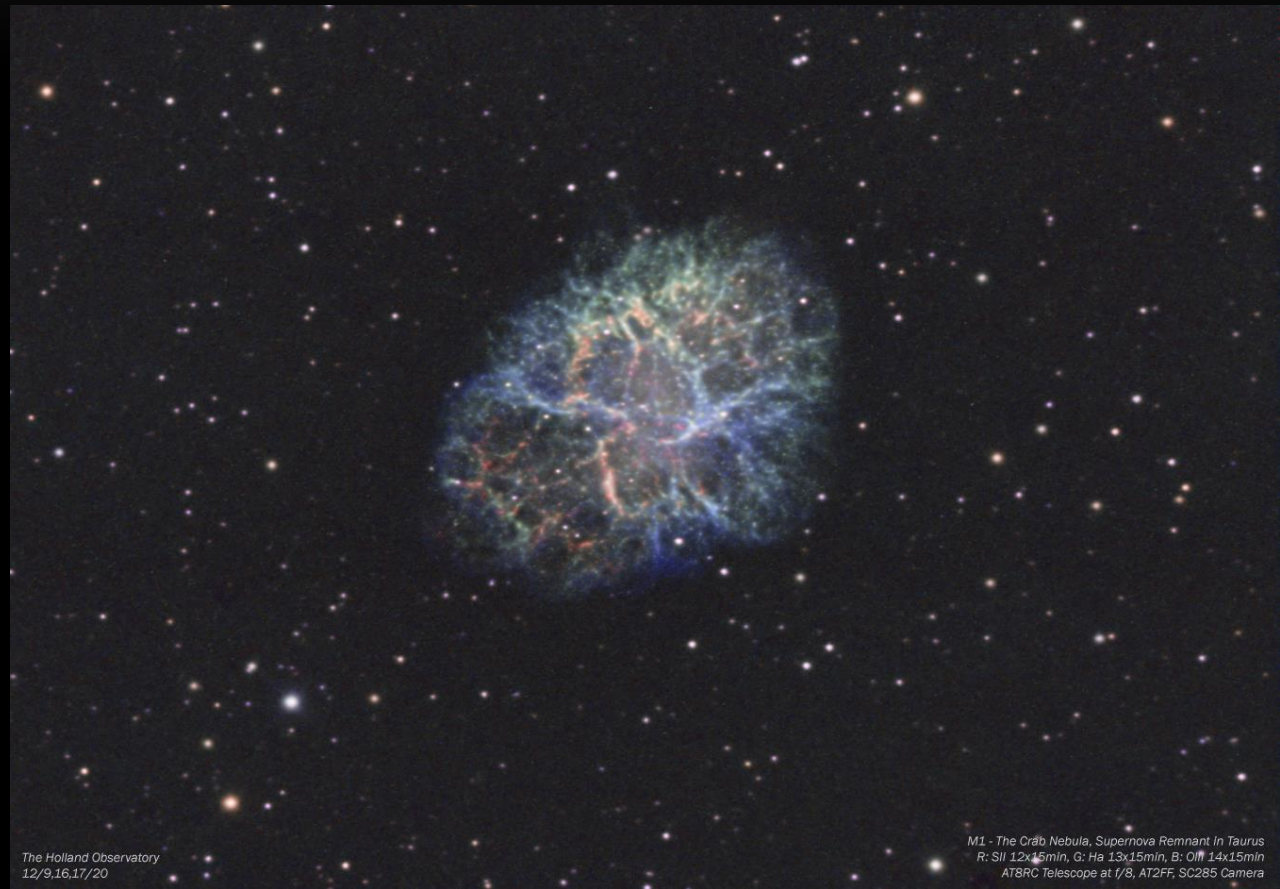
The Holland Observatory
10/15/14 & 11/18/14

Melotte 15 / IC1805 Heart of the Heart Nebula in Cassiopeia
Narrowband - Ha: 14x12min, SII: 15x12min, OIII: 11x12min
200mm Newtonian at f/5, MPCC, SC285 Camera

What else can we image with Narrowband Filters?

1. Supernova Remnants
2. Planetary Nebulae
3. Wolf-Rayet Stars

1. Supernova Remnants



2. Planetary Nebulae



The Holland Observatory
9/25,26/11

NGC7293, The Helix Nebula in Aquarius
R(Ha), G(SII), B(OIII); Ha 11x5m, SII 7x4m, OIII 11x5m
200mm Newtonian at f/5, SC285 Camera

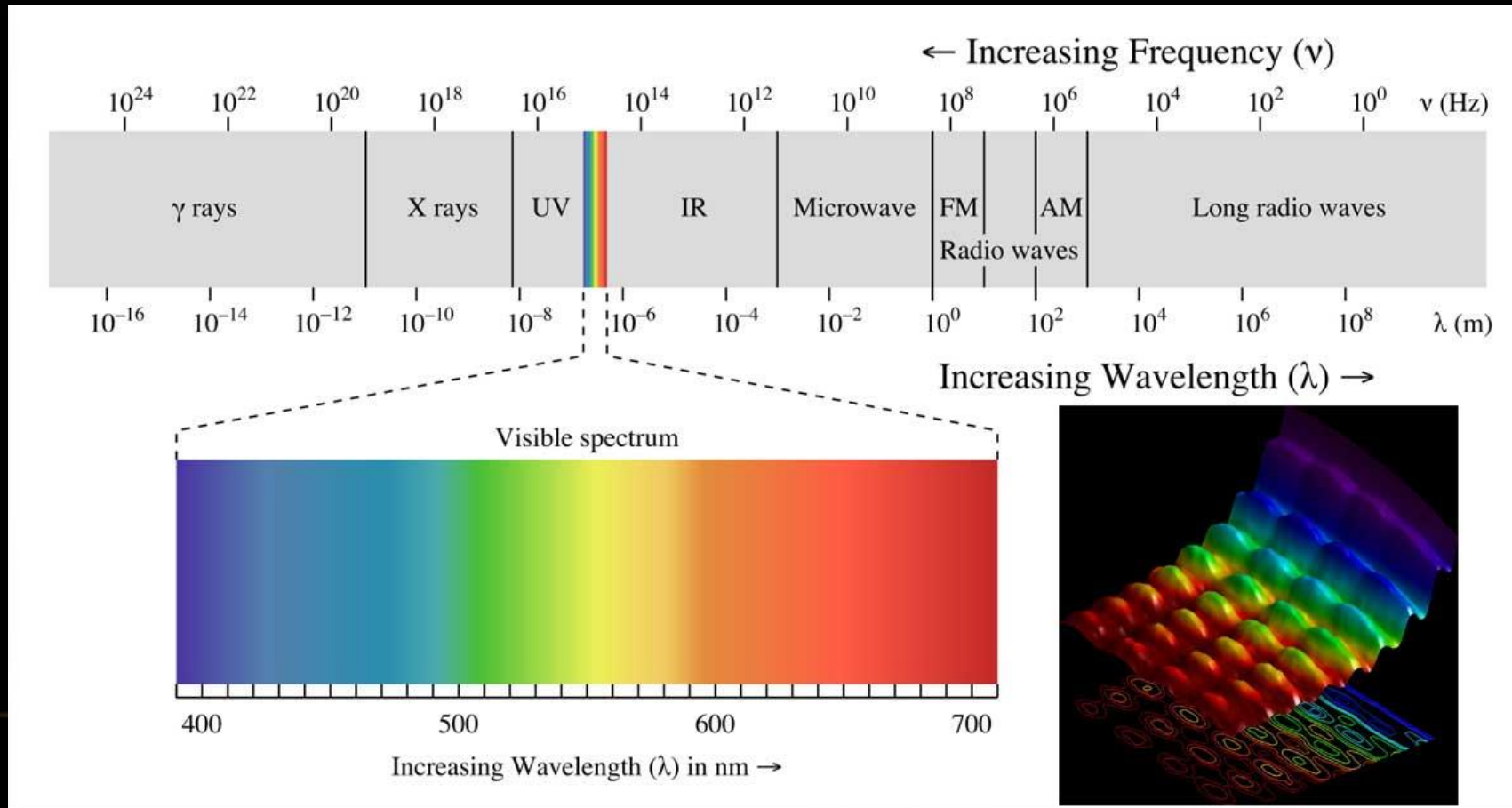
3. Wolf-Rayet Stars (First noted in 1867 by French Astronomers Charles Wolf and George Rayet, the spectra of Wolf-Rayet stars are characterized by broad band emission lines predominantly of helium. Massive, luminous, and hot (likely descended from stars of spectral class O), they develop intense radiation pressure and winds. A Wolf-Rayet's violent behavior results in tremendous mass-loss, making the star unstable. It's hydrogen-rich atmosphere is stripped away. Intense radiation from the exposed helium core expels gas at high-velocity that collides into a previously ejected envelope to create the shell.)



What else can we do?

What about other types of filters?

It turns out that most light pollution occurs in Visible portion of spectrum

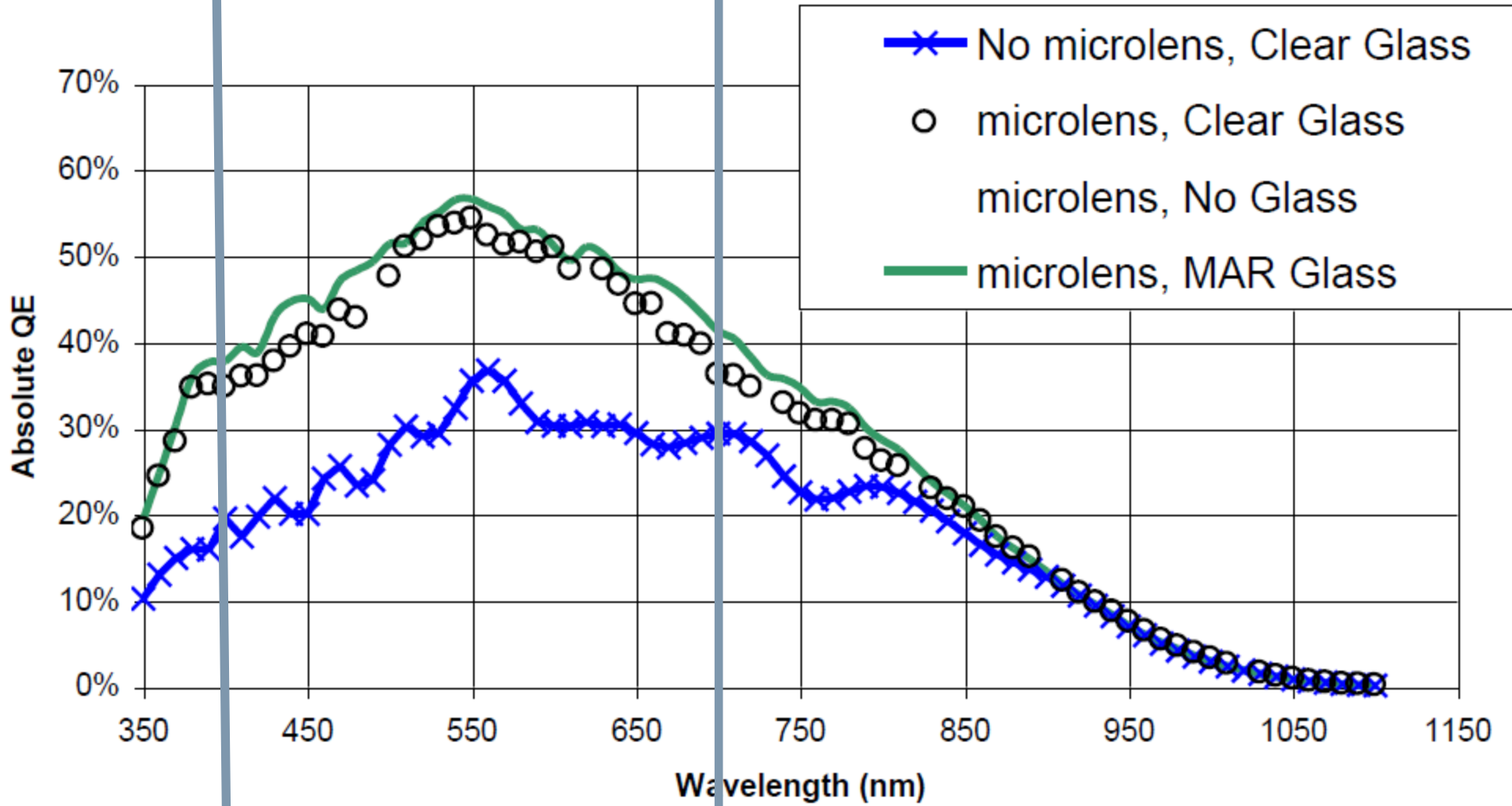


Ultraviolet

Visible Spectrum

Near Infrared (NIR)

KAF-8300 Quantum Efficiency



NIR displays
different
features than
visible range



NIR makes
imaging galaxies
possible from
light polluted
areas

The Holland Observatory
1/17/13 and 2/13, 14, 16/13

M51 - The Whirlpool Galaxy in Canes Venatici
NIR - L: NIRL 18x10min; R: >880nm 15x10min + 3x10min; G: 800-900nm 12x10min; B: 700-800nm 11x10min
200mm Newtonian f/5, SC285 Camera

What have we done so far?

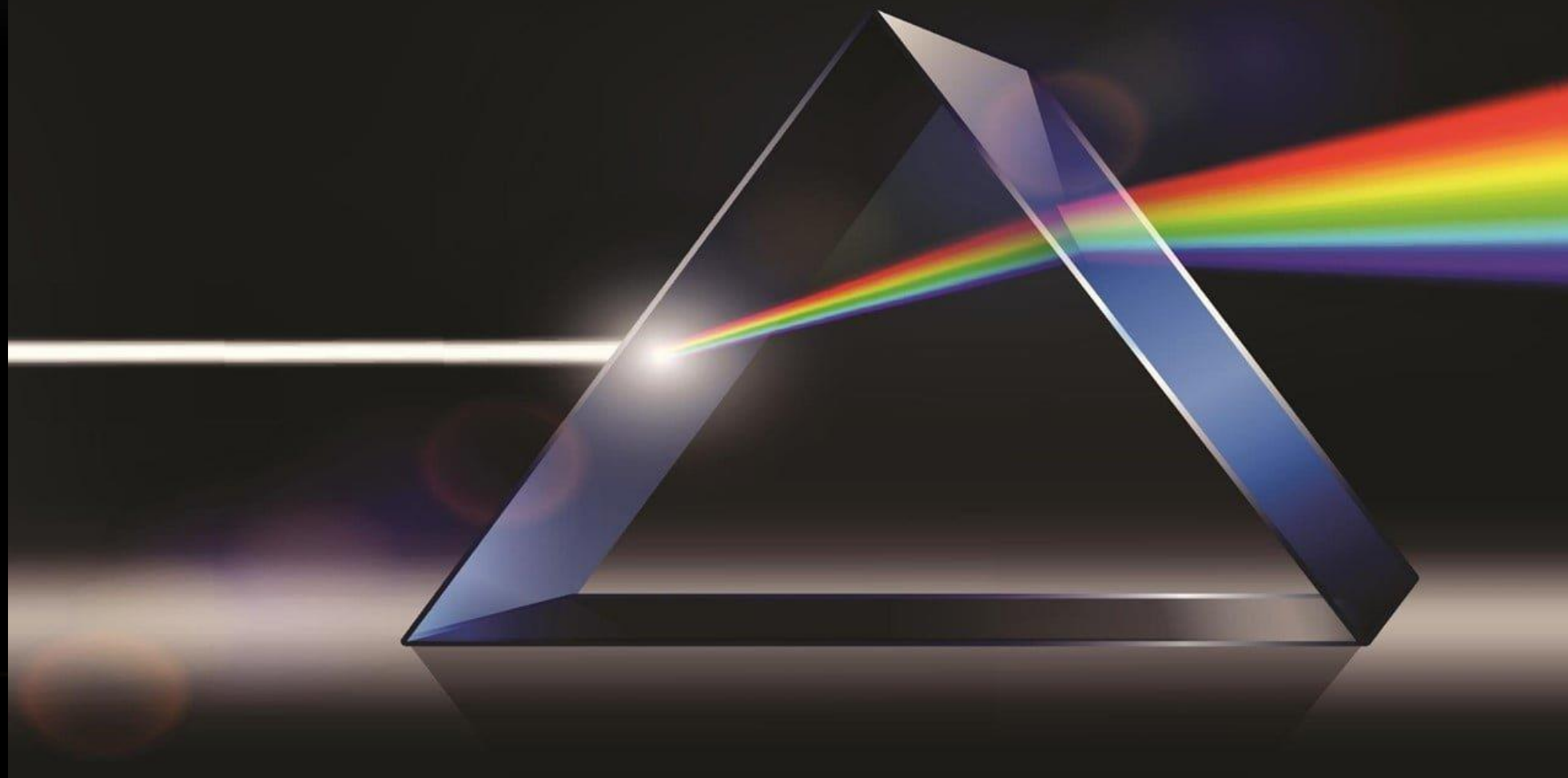
- Solar System: Sun, Moon, planets
- Double Stars, Globular Clusters, Open Clusters
- Emission Nebula – Broadband and Narrowband
- Supernova Nebula, Planetary Nebula, Wolf-Rayet Stars - Narrowband
- NIR Imaging

What else can we do?

How about something that is not imaging?

How about something more Sciencey?

Spectroscopy

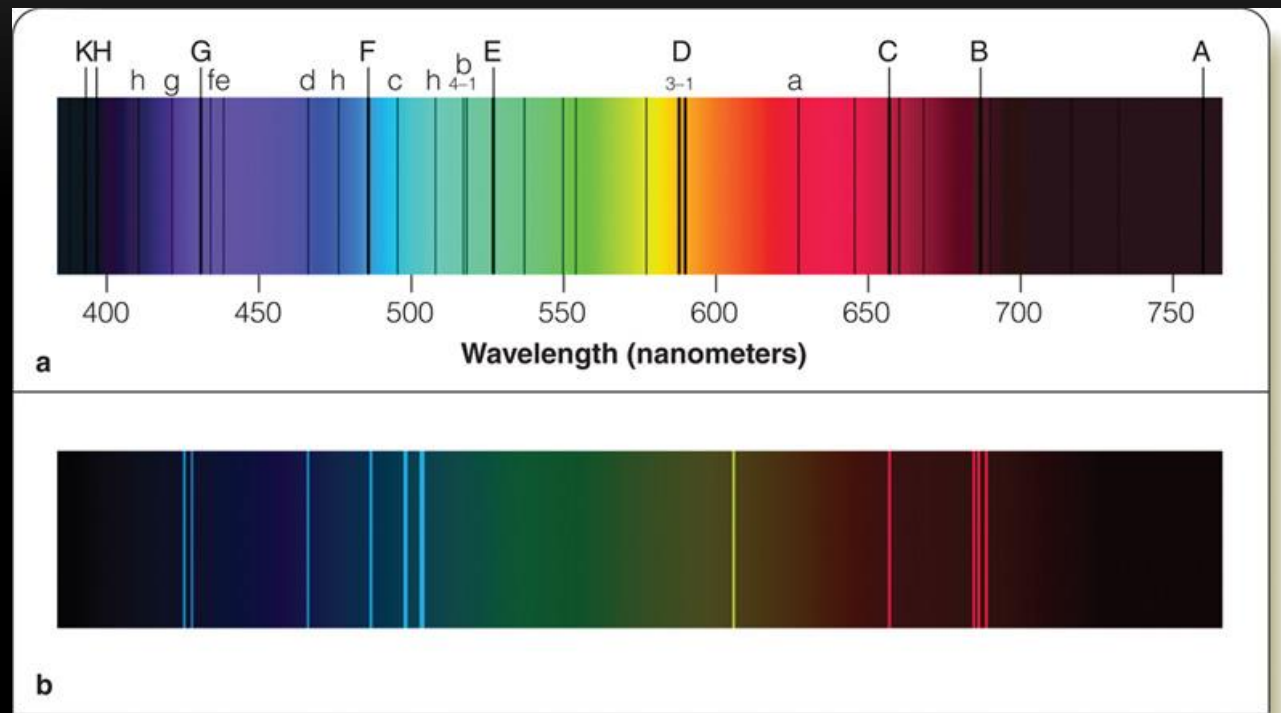


Spectrum of the Sun



Project STAR Spectrometer, \$36
(Flame spectra, street lights, solar)

The Sun's Spectrum



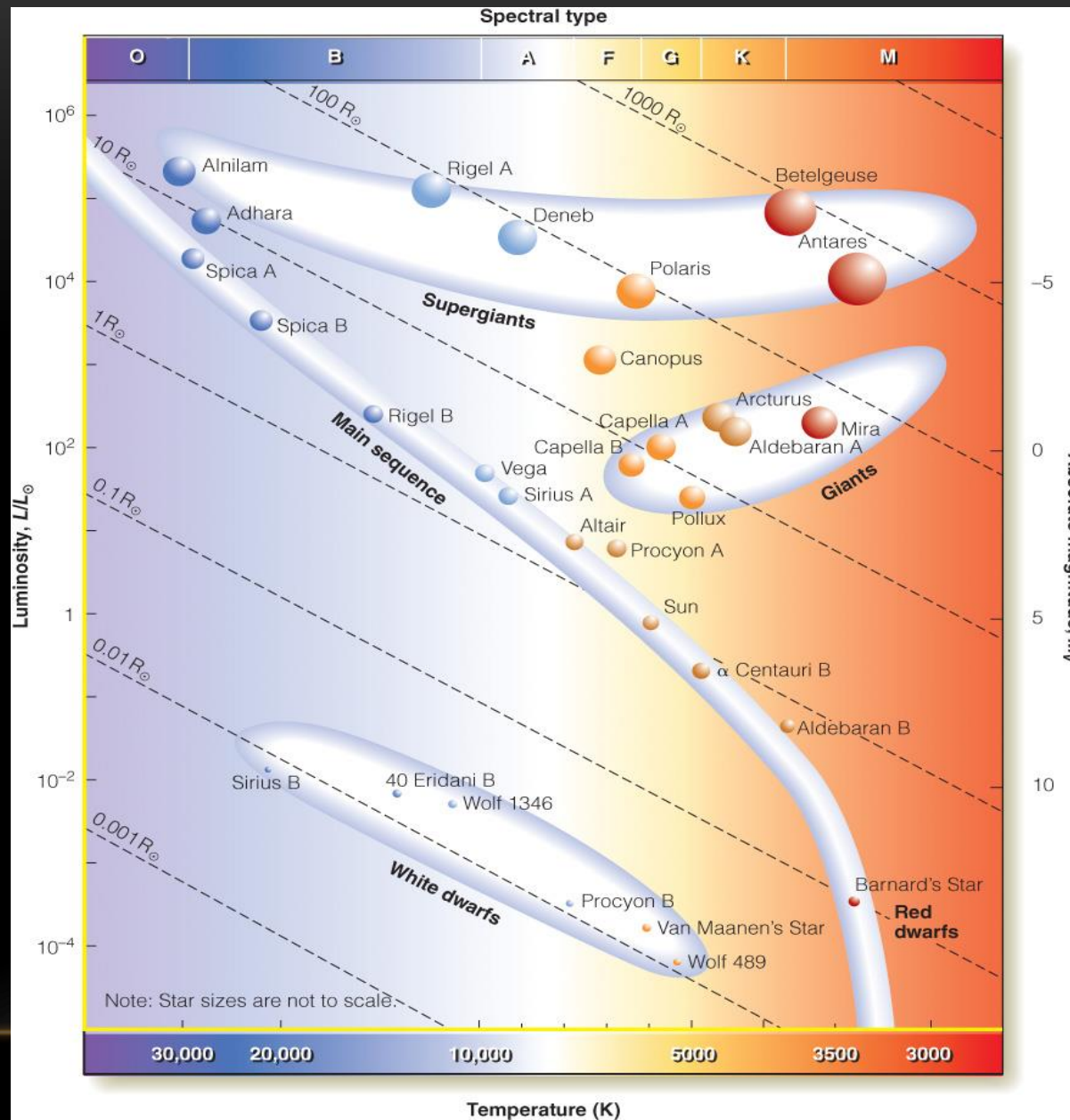
Foundations of Astronomy (Seeds & Backman)

Spectrum of Stars & Deep Sky Objects

Field Tested Systems - produces hardware & software which enables spectroscopy

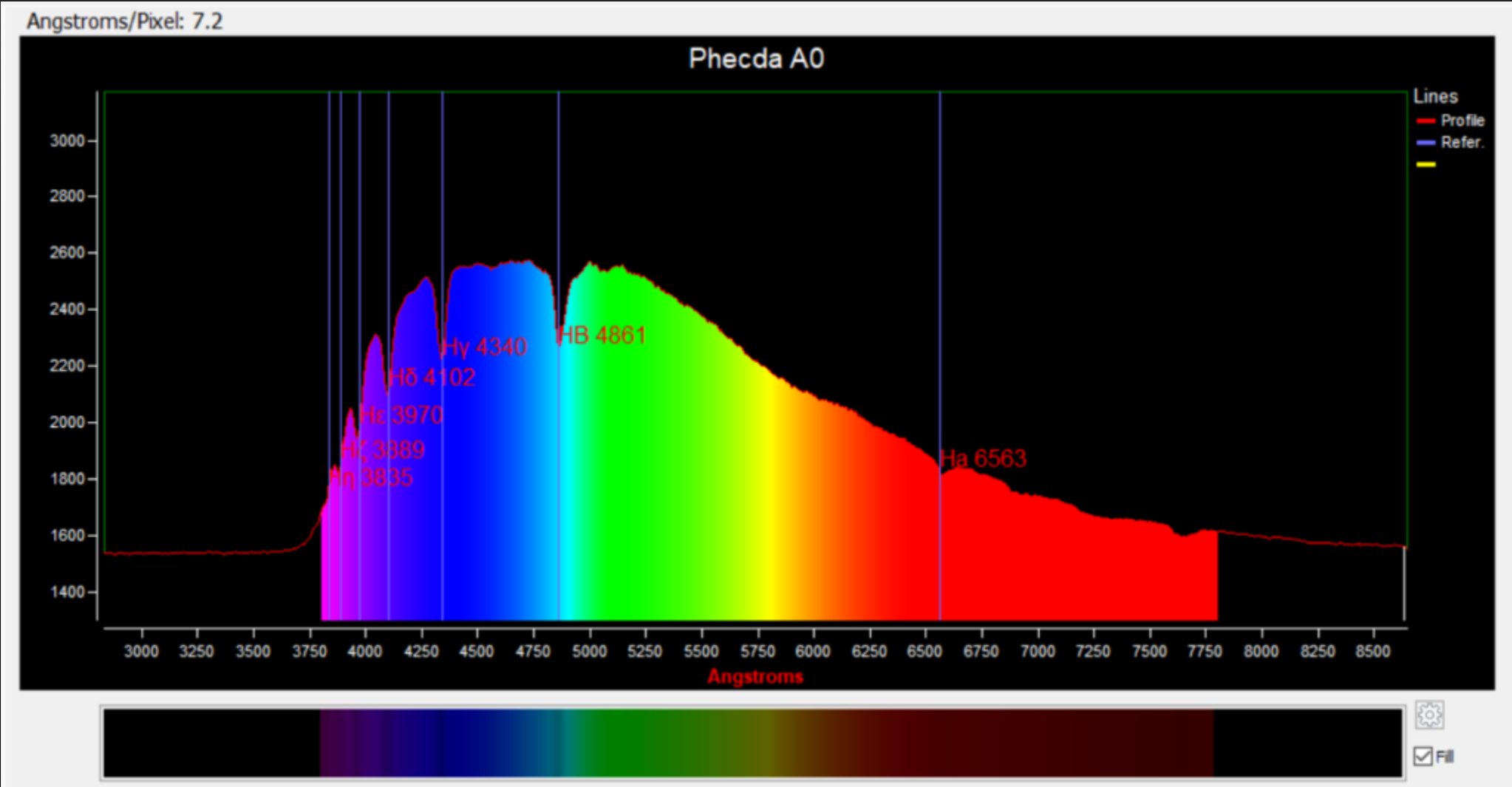
Rspec (Real-time Spectroscopy) at <https://www.rspec-astro.com/>

Hertzsprung-Russell (H-R) Diagram



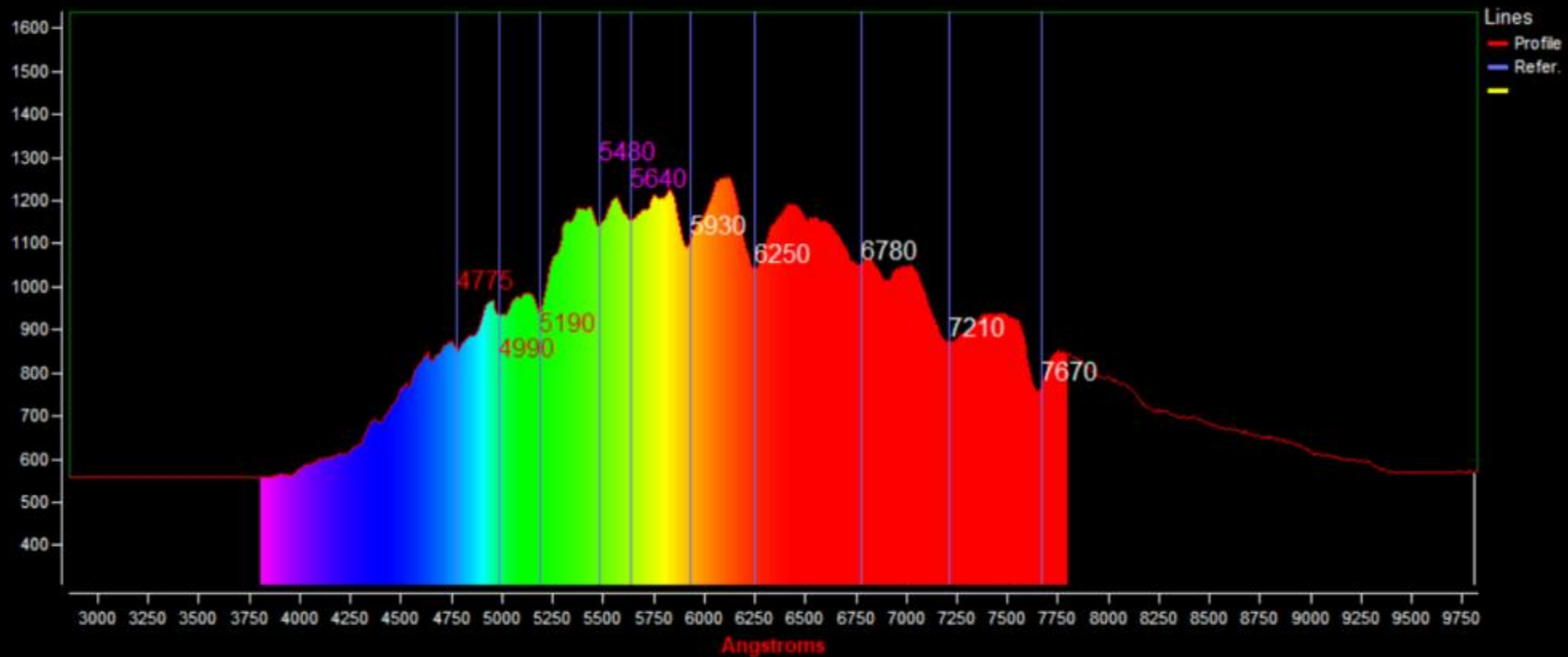
(Foundations of Astronomy – Seeds & Backman)

Stellar Spectra



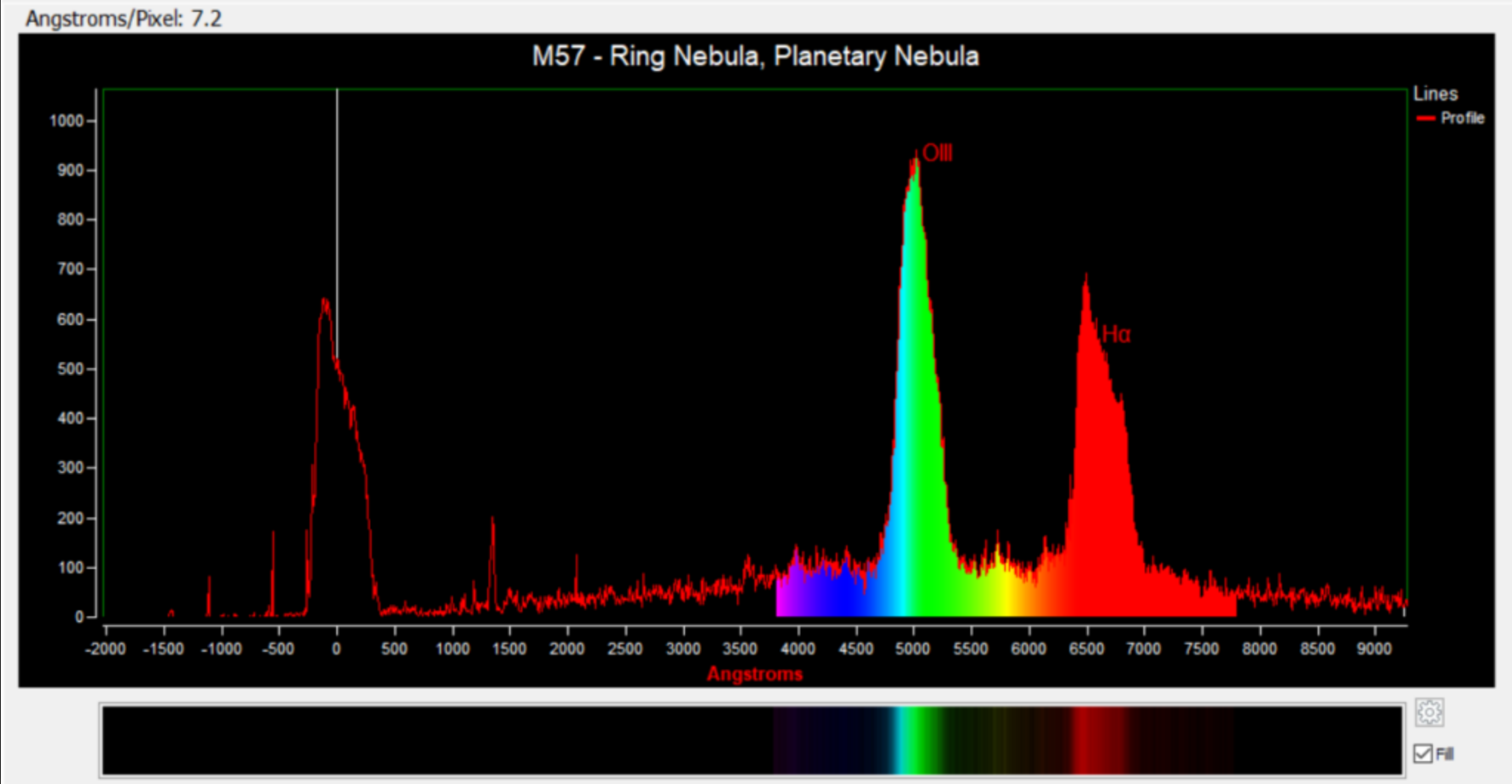
Angstroms/Pixel: 7.2

Yed Prior M1 - Titanium Oxide (TiO)



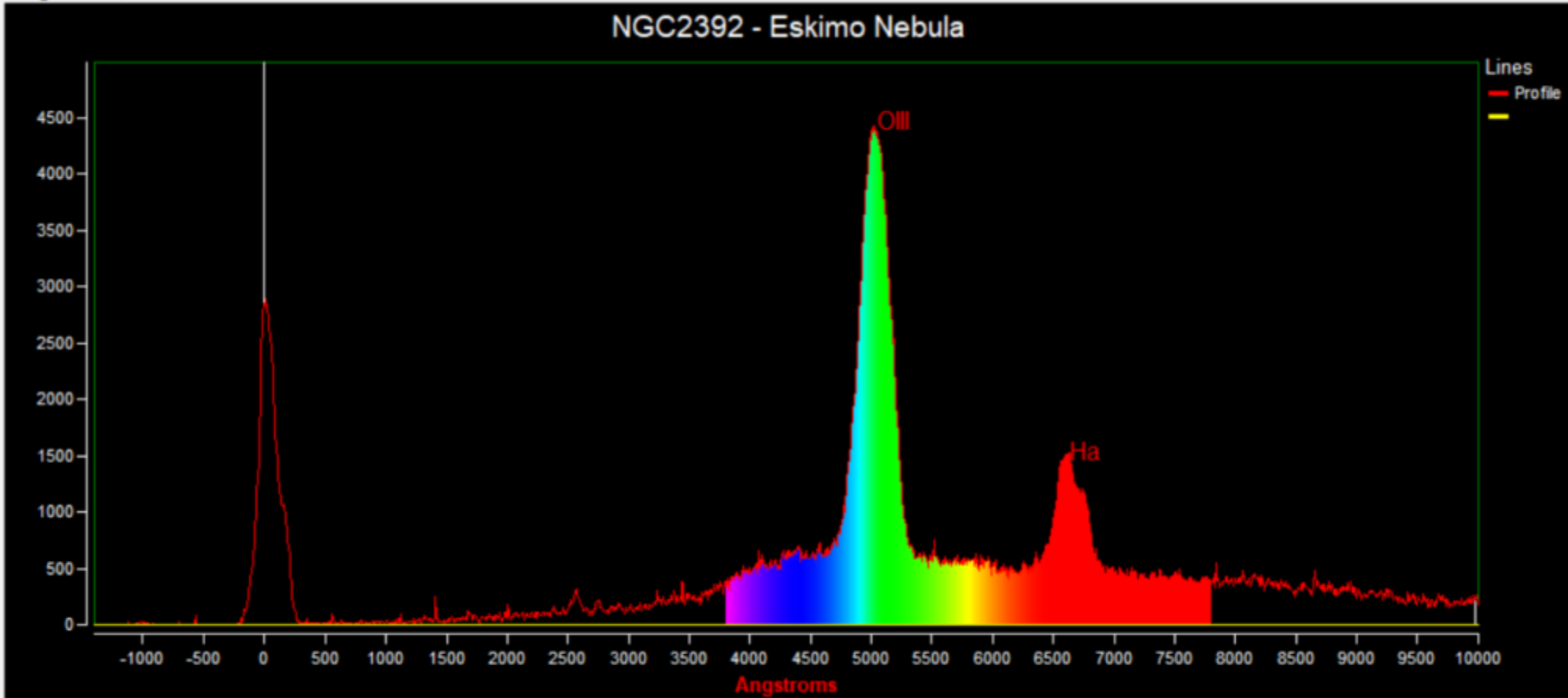
Settings icon (gear) and a checked "Fill" checkbox.

Planetary Nebula Spectra



Angstroms/Pixel: 7.2

NGC2392 - Eskimo Nebula



Settings icons: a gear icon and a checked checkbox labeled 'Fill'.

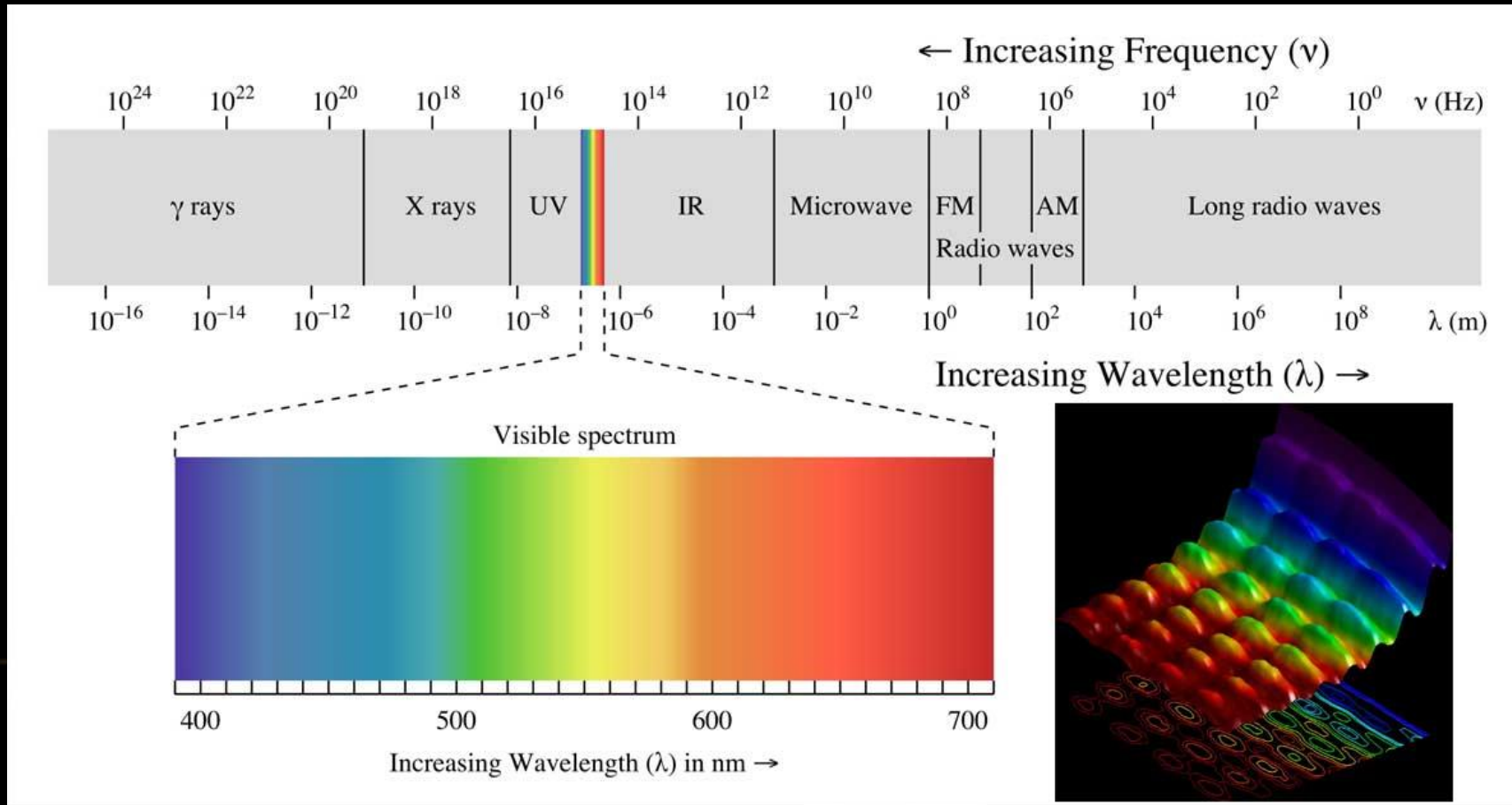
What else can we do?

One more thing =>



Radio Astronomy

- Can be done day or night
- Light pollution plays no role
- Can be done with or without clouds





21cm Radio Telescope
(21cm => 1.42GHz)
Neutral Hydrogen Emission

WiFi Dish Antenna:

eBay

- 24dBi gain
- 2.4GHz but works at 1.42GHz
- 14°x10° beam width

Low Noise Amplifier:

Amazon

- 40dB, Tuned for 21cm

Software Defined Radio:

<https://www.rtl-sdr.com/>

- USB interface

SDR# software -



21cm Radio Telescope
(21cm => 1.42GHz)
Hydrogen Emission

- Raspberry Pi Configuration



Sawbird Low Noise Amplifier
with Hydrogen bandpass
(1420MHz)

Software Defined Radio
(SDR) receiver

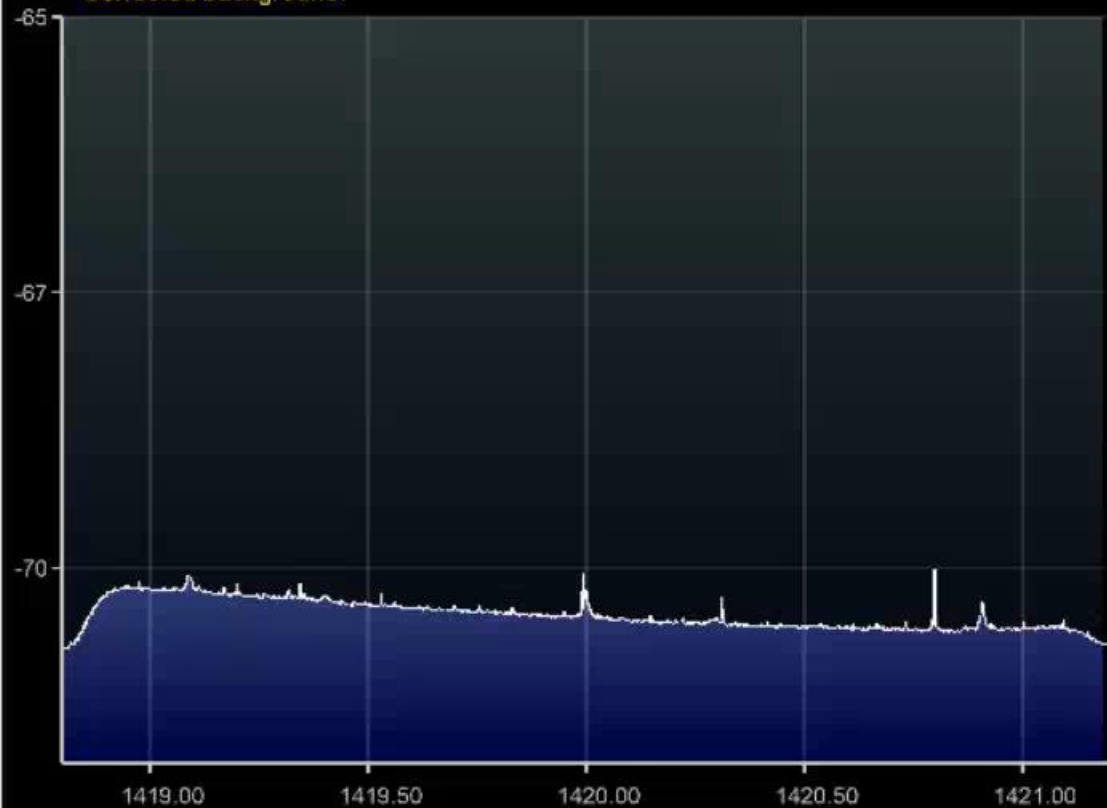
Raspberry Pi 3B+

Project #1 – Create video of one day of Milky Way passing over

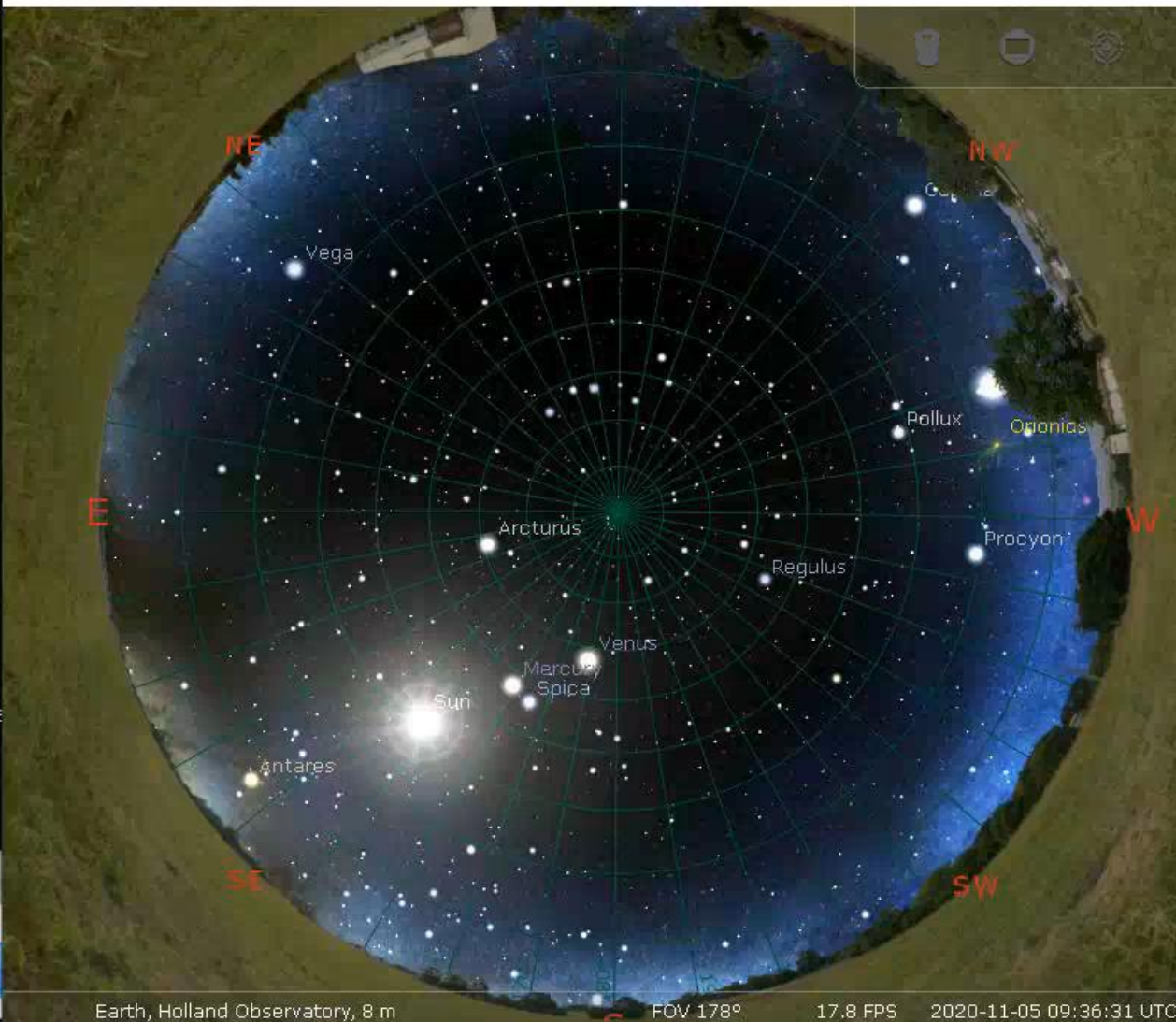
IF Average Window

1418.6623327 MHz Cumulations: 87000 of: 902000 Recording time: 37s

Corrected background!



Stellarium 0.20.3



AIRSPY SDR# v1.0.0.1769 - Spy Server Network

Control panel for the SDR software, featuring icons for menu, stop, settings, mute, and a frequency slider.

001.420.000.000

- ▶ IF Noise Blanker *
- ▶ Demodulator Noise Blanker *
- ▶ Recording *
- ▶ Zoom FFT *

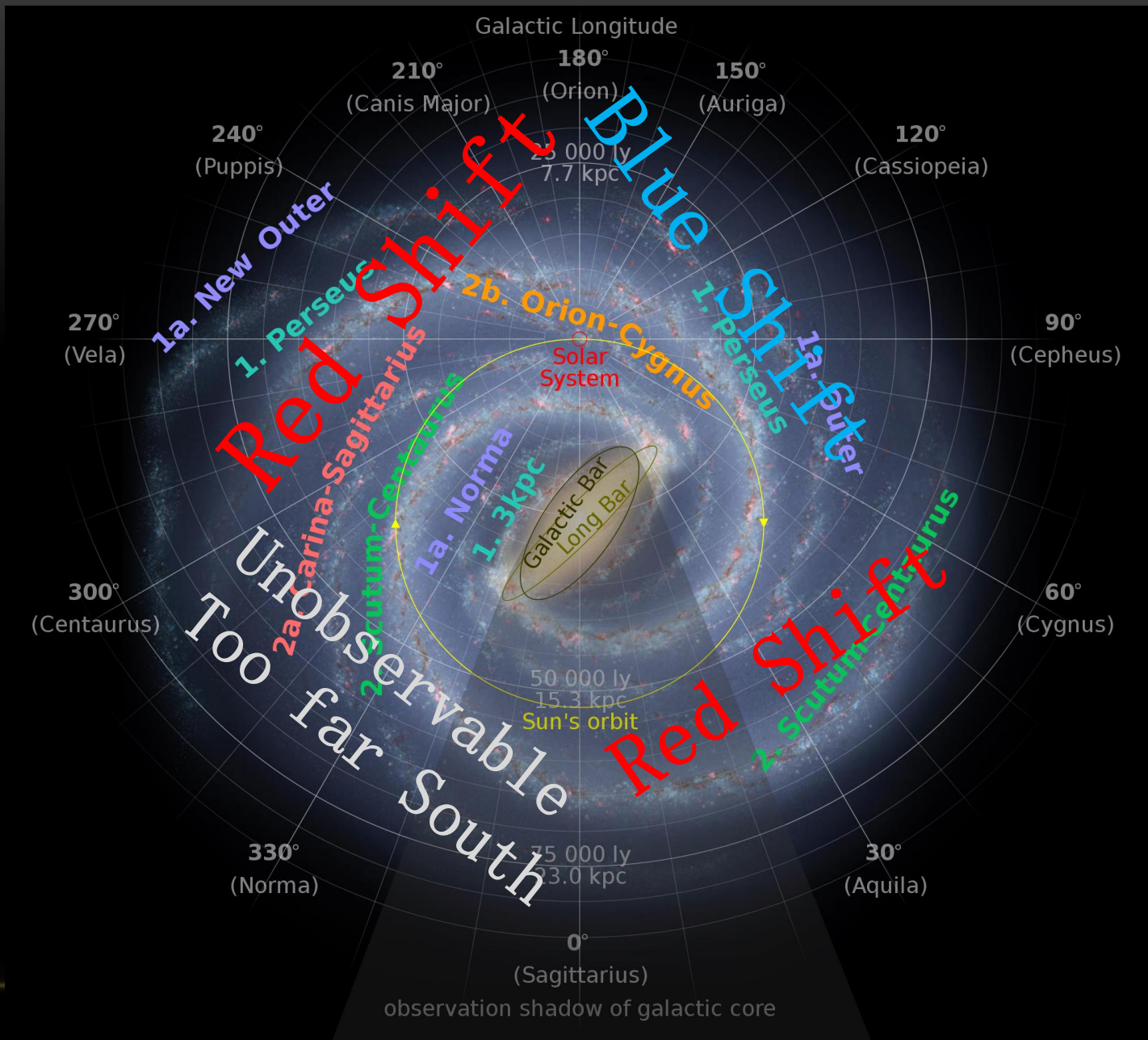
dBFS
0
-10
-20
-30

2020-11-05 09:36:31.35

Earth, Holland Observatory, 8 m FOV 178° 17.8 FPS 2020-11-05 09:36:31 UTC

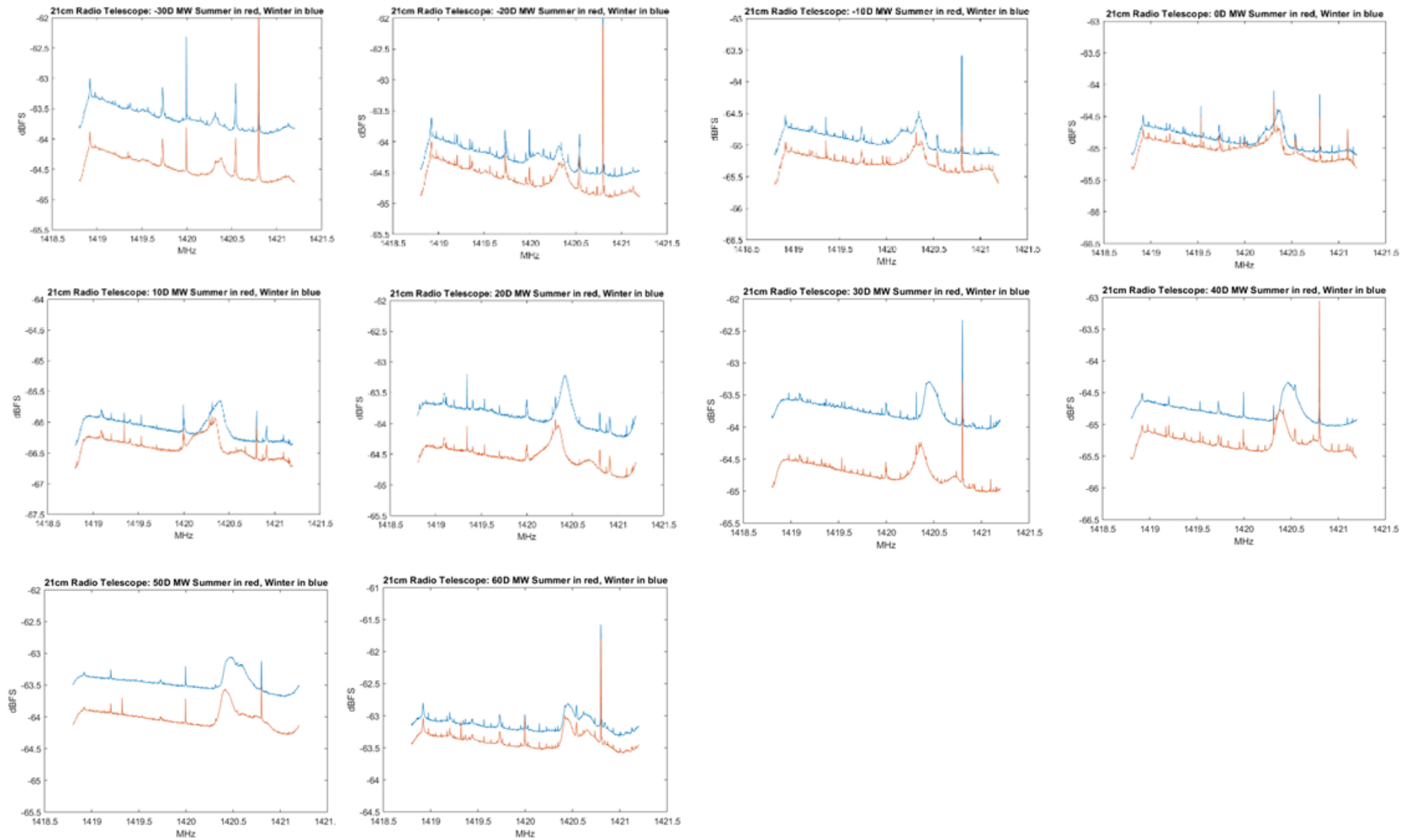
Zoom control panel with a vertical slider and a 'Zoom' label.

Project #2 – Measure relative location of our Solar System within Milky Way



<https://earthsky.org/upl/2020/01/milky-way-arms-suns-location-orion-cygnus-arm.png>

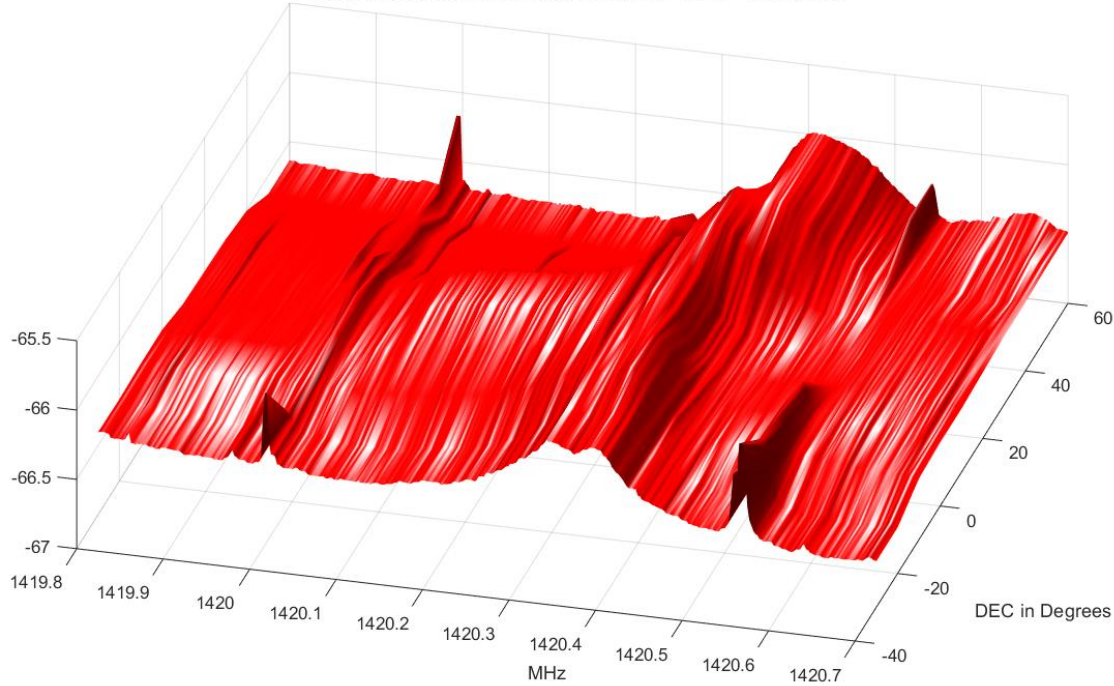
21cm Plots at Galactic Equator – Spaced at 10 Degrees Declination



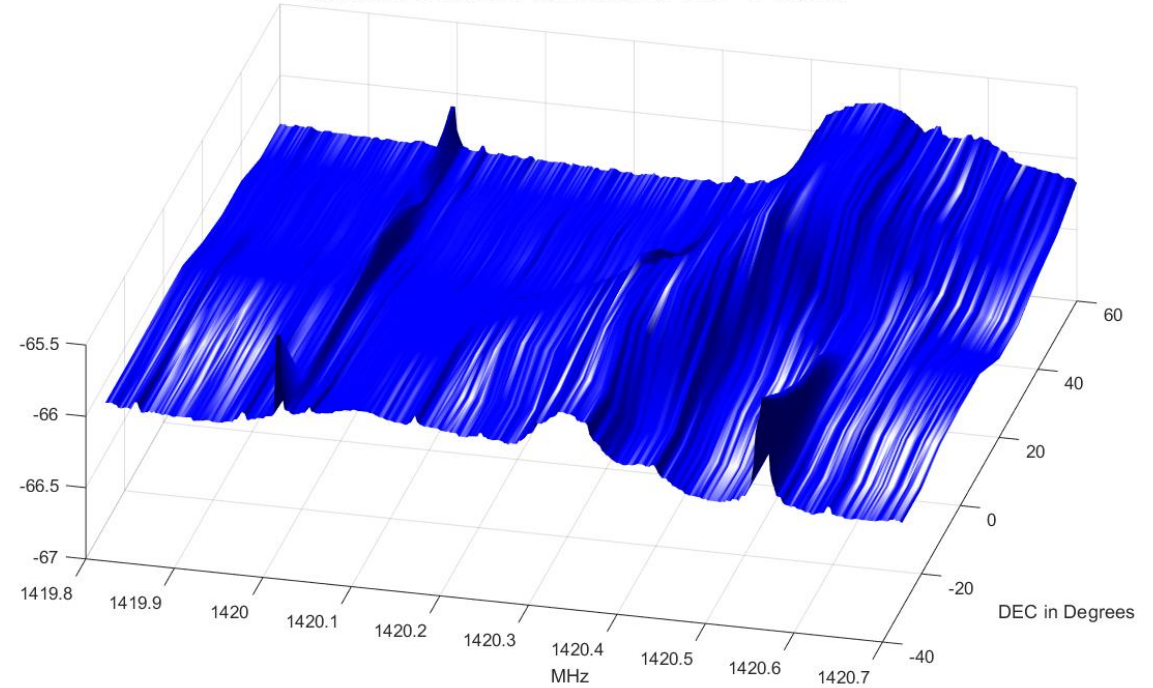
3D Surface Plots of Summer (Red) and Winter (Blue) Milk Way Data

- The **diameter** of the luminous **Milky Way** is between 100,000 and 120,000 light years across.
- Sun (Solar System) is 1/2 to 2/3rds from center (www.universetoday.com)

21cm Radio Telescope: Summer MW DEC -30 to +60 Degrees



21cm Radio Telescope: Winter MW DEC -30 to +60 Degrees

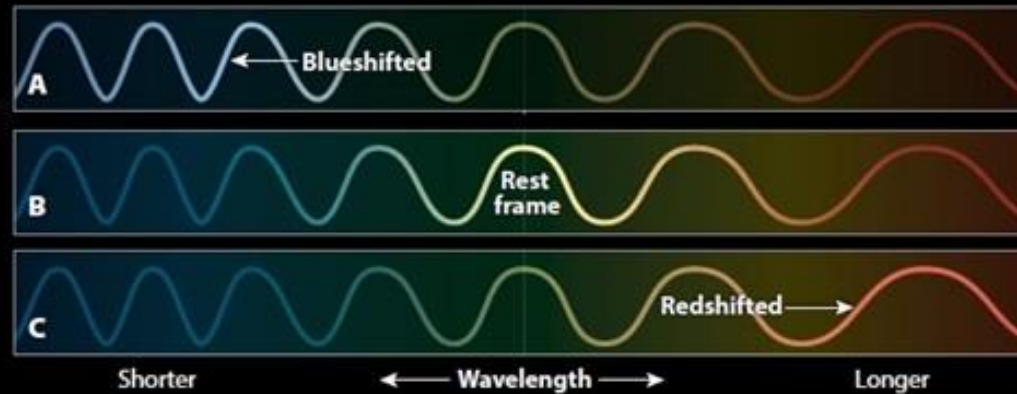
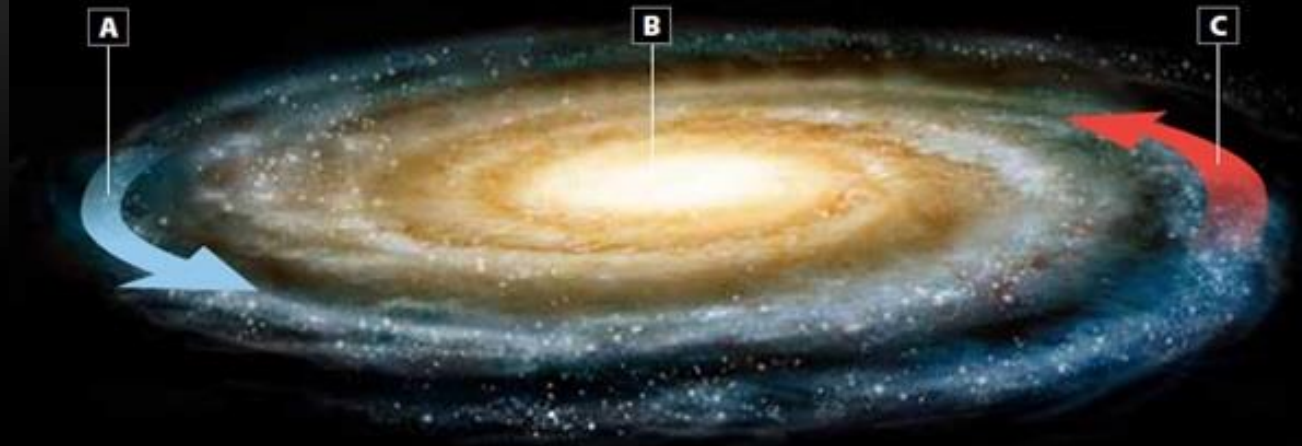


Declination vs. Amplitude vs. Frequency
-30° DEC: Center of galaxy (Sagittarius)
+60° DEC: Most Northerly point of galactic equator (all data taken along galactic equator)

=> Transition between red shift and blue shift is approx. 1/2 to 2/3 from center of galaxy to most Northerly point

Project #3 – Measure the rotation of the Milky Way

Measuring a galaxy's rotation



As a galaxy rotates, the material moving away from us shows a redshift in the wavelength of any emitted light (red arrow). Material moving toward us shows a blueshift (blue arrow). By measuring these shifts across a galaxy, astronomers can determine its rotation. ASTRONOMY: ROEN KELLY

<https://astronomy.com/magazine/ask-astro/2018/05/rotational-speed-of-a-galaxy>

To Summarize => What Astronomy Can We Do from a Light Polluted Area?

- Solar System: Sun, Moon, planets
- Double Stars, Globular Clusters, Open Clusters
- Emission Nebula – Broadband and Narrowband
- Supernova Remnant, Planetary Nebula, Wolf-Rayet Stars - Narrowband
- NIR Imaging
- Spectroscopy
- Radio Astronomy



The End

www.holland-observatory.net
